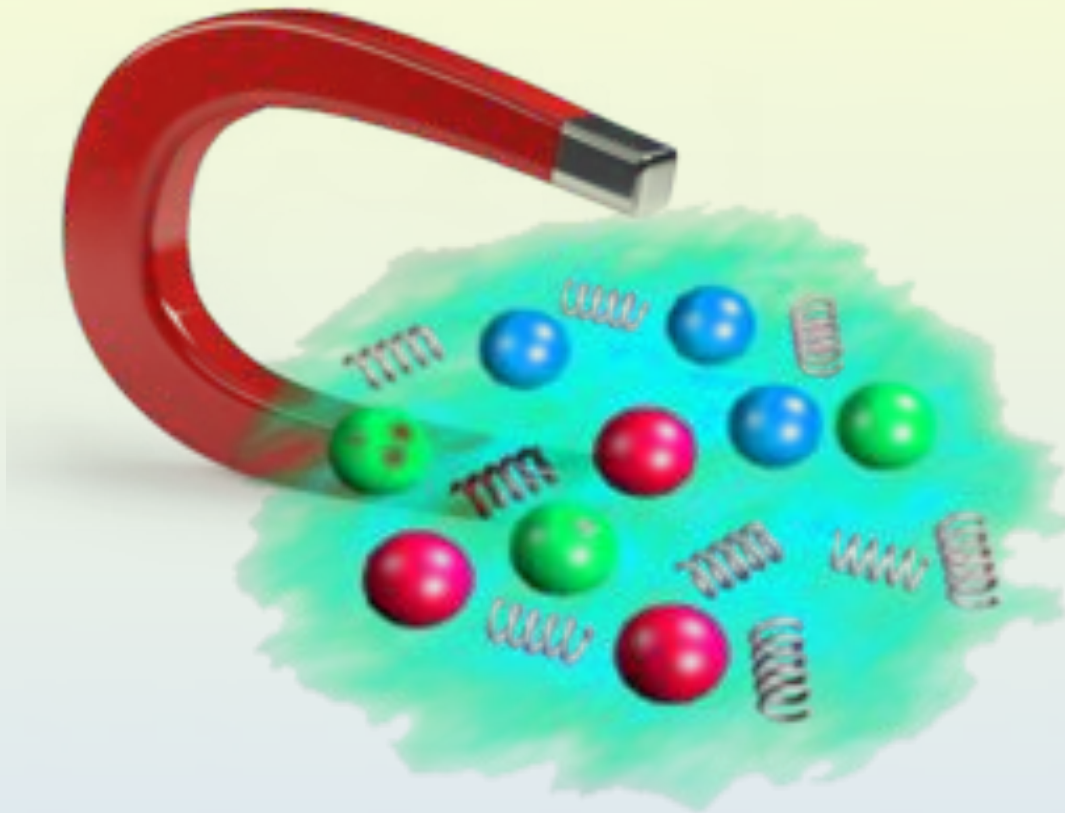


# Magnetized Relativistic Matter: Generalized Landau-level representation & its applications

**Igor Shovkovy**

**Arizona State University**

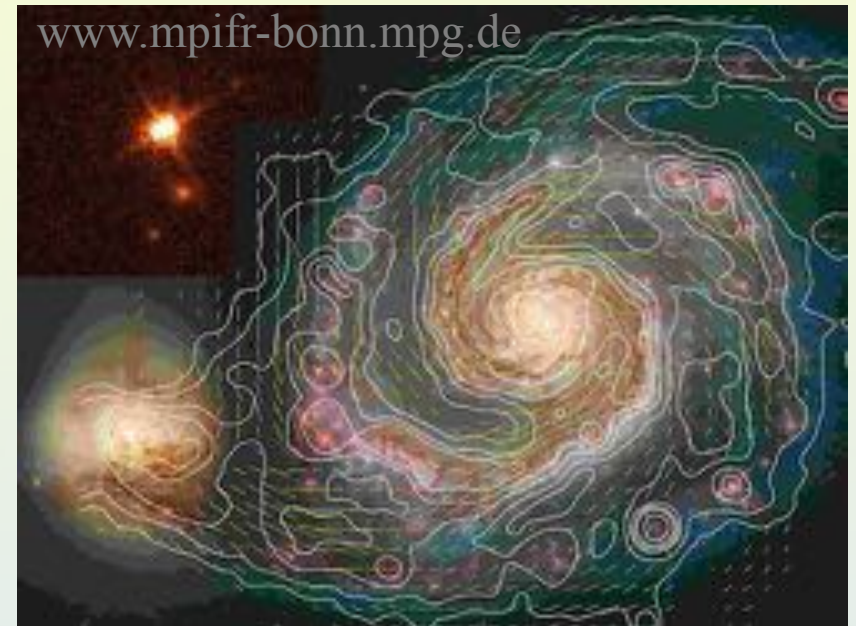




# MAGNETIC FIELDS EVERYWHERE

[Miransky & Shovkovy, Physics Reports 576 (2015) pp. 1-209]

- Current galactic magnetic fields  $\sim 10^{-6}$  G
- Current magnetic fields in voids  $\sim 10^{-15}$  G
- Problem of magnetogenesis in Early Universe
- Perhaps during the electro-weak phase transition  
–  $10^{20}$  to  $10^{24}$  G ( $\sim 1$  GeV to 100 GeV)

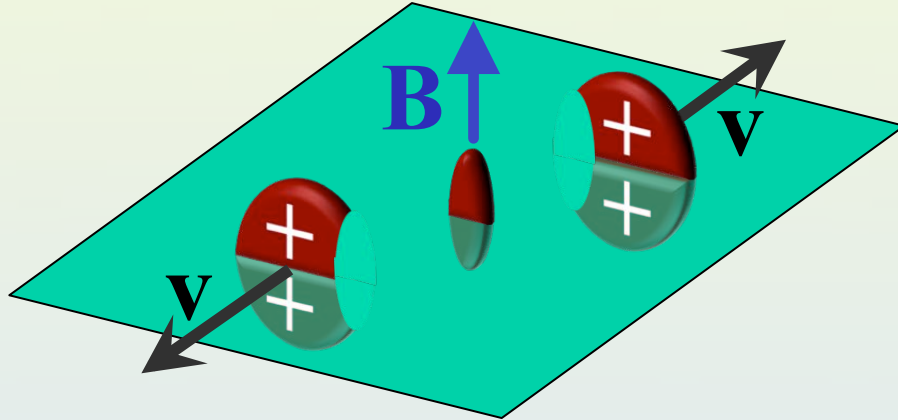


- Magnetized dense baryonic matter
  - $10^{10}$  to  $10^{18}$  G (10 keV to 100 MeV)
- Magnetic field may affect
  - Competition of ground state phases
  - EoS of dense baryonic matter
  - the M-R relation of compact stars
  - Transport and emission properties
  - Evolution of supernovas & protoneutron stars



# Little Bangs

- Magnetized QGP at RHIC/LHC
  - $10^{18}$  to  $10^{19}$  G ( $\sim 100$  MeV)



[Rafelski & Müller, PRL, 36, 517 (1976)],  
 [Kharzeev et al., arXiv:0711.0950],  
 [Skokov et al., arXiv:0907.1396],  
 [Voronyuk et al., arXiv:1103.4239],  
 [Bzdak & Skokov, arXiv:1111.1949],  
 [Deng & Huang, arXiv:1201.5108]

- Using Lienard-Wiechert potentials,

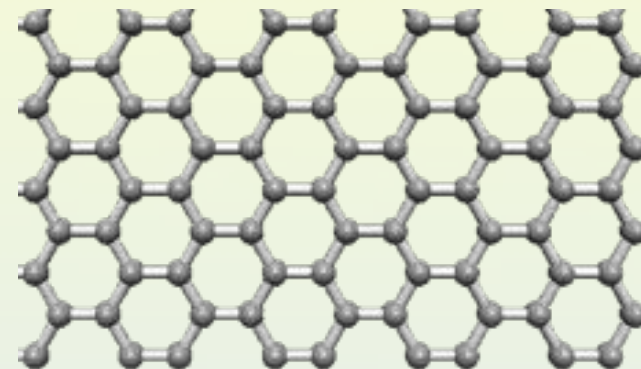
$$e\mathbf{E}(t, \mathbf{r}) = \frac{e^2}{4\pi} \sum_n Z_n \frac{\mathbf{R}_n - R_n \mathbf{v}_n}{(R_n - \mathbf{R}_n \cdot \mathbf{v}_n)^3} (1 - v_n^2)$$

$$e\mathbf{B}(t, \mathbf{r}) = \frac{e^2}{4\pi} \sum_n Z_n \frac{\mathbf{v}_n \times \mathbf{R}_n}{(R_n - \mathbf{R}_n \cdot \mathbf{v}_n)^3} (1 - v_n^2)$$

# Dirac/Weyl materials

- High magnetic field lab
  - $10^5$  G ( $\sim 100$  meV @  $v_F=c/300$ )

- Graphene



- 3D materials with Dirac/Weyl quasiparticles

- $\text{Bi}_{1-x}\text{Sb}_x$  alloy (at  $x \approx 4\%$ )
- $\text{Na}_3\text{Bi}$
- $\text{Cd}_3\text{As}_2$
- $\text{ZrTe}_5$
- TaAs, NbAs, TaP, ...

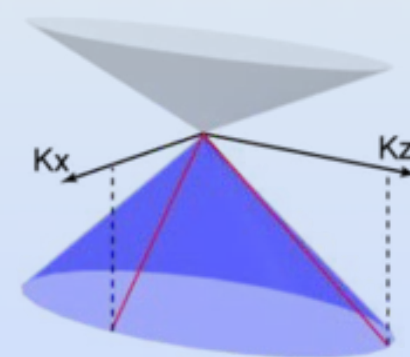
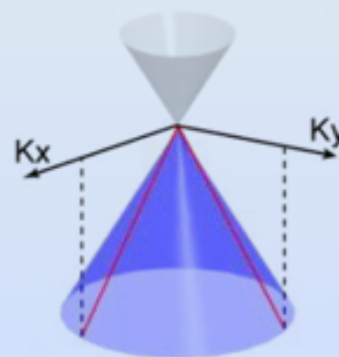
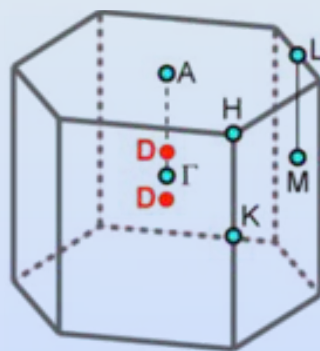
[Z. K. Liu et al., arXiv:1310.0391]

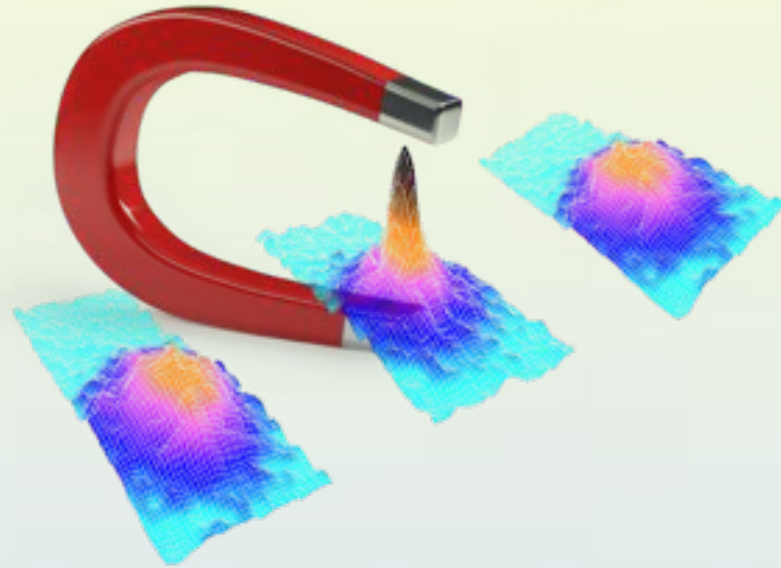
[M. Neupane et al., arXiv:1309.7892]

[S. Borisenko et al., arXiv:1309.7978]

[X. Li et al., arXiv:1412.6543]

[arXiv:1502.03807, arXiv:1502.04684,  
arXiv:1504.01350, arXiv:1507.00521]





# RESEARCH DIRECTIONS

[Miransky & Shovkovy, Physics Reports **576** (2015) pp. 1-209]

# Magnetic catalysis

- Magnetic catalysis of chiral symmetry breaking & anisotropic confinement (vacuum QCD, QCD @  $T \neq 0$ )

$$E_n^{(3+1)}(p_3) = \pm \sqrt{2n|eB| + p_3^2} \quad \left. \vphantom{E_n^{(3+1)}(p_3)} \right\} \Rightarrow m_{\text{dyn}}(B) \neq 0$$

$(D \rightarrow D - 2)$

- $T=0$ : catalysis

[Miransky & I.S., Phys. Rev. D **66** (2002) 045006]

[I.S., Lect. Notes Phys. **871**, 13 (2013)]

[Bali et al., Phys. Rev. D **86**, 071502 (2012)]

...

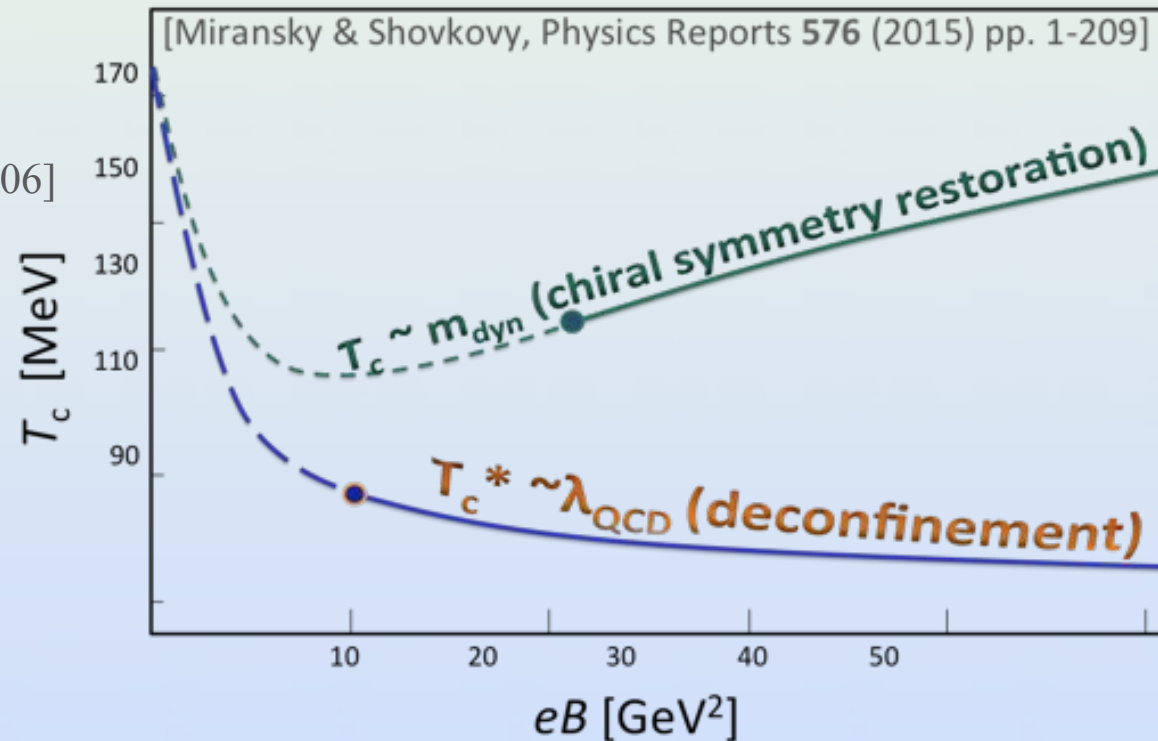
- $T_c(B)$ : inverse catalysis

[Bali et al., JHEP **02**, 044 (2012)]

[Bali et al., PRD **86**, 071502 (2012)]

[G. Endrodi, arXiv:1504.08280]

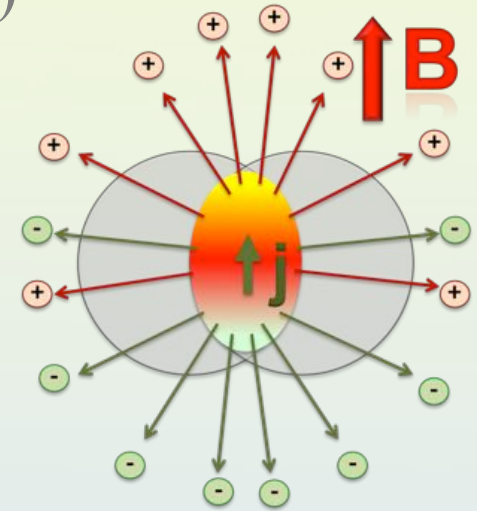
...





- Chiral magnetic/separation effects, chiral magnetic waves (correlations of charged particle in HIC)

$$\langle \vec{j} \rangle = \frac{e^2 \vec{B}}{2\pi^2} \mu_5$$



- Signs of local P-violation?

$$\frac{d(N_R - N_L)}{dt} = -\frac{g^2 N_f}{16\pi^2} \int d^3x F_a^{\mu\nu} \tilde{F}_{\mu\nu}^a$$

[Kharzeev, McLerran, Warringa, Nucl. Phys. A **803**, 227 (2008)]

[Fukushima, Kharzeev, Warringa, Phys. Rev. D **78**, 074033 (2008)]

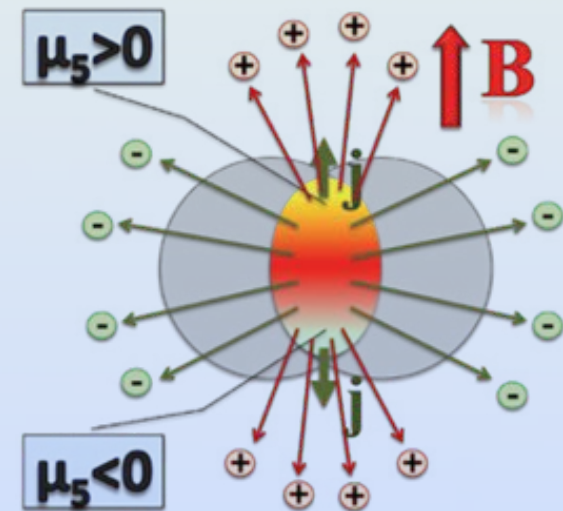
...

- Signs of a chiral magnetic wave?

[Yee, Kharzeev, Phys. Rev. D **83**, 085007 (2011)]

[Gorbar, Miransky, Shovkovy, Phys. Rev. D **83**, 085003 (2011)]

[Burnier, Kharzeev, Liao, Yee, Phys. Rev. Lett. **107** (2011) 052303]



# Magnetic helicity

- Magnetic helicity evolution in the Early Universe & in HIC (magnetogenesis,  $\mathcal{H}_k \rightarrow \mu_5 \rightarrow \mathcal{H}_k$ )

$$\frac{d(n_L - n_R)}{dt} = \frac{2\alpha}{\pi} \frac{1}{V} \int d^3x (\vec{E} \cdot \vec{B})$$

$$\nabla \times \vec{B} = \sigma \vec{E} + \frac{\alpha}{\pi} \mu_5 \vec{B} + \frac{\partial \vec{E}}{\partial t}$$

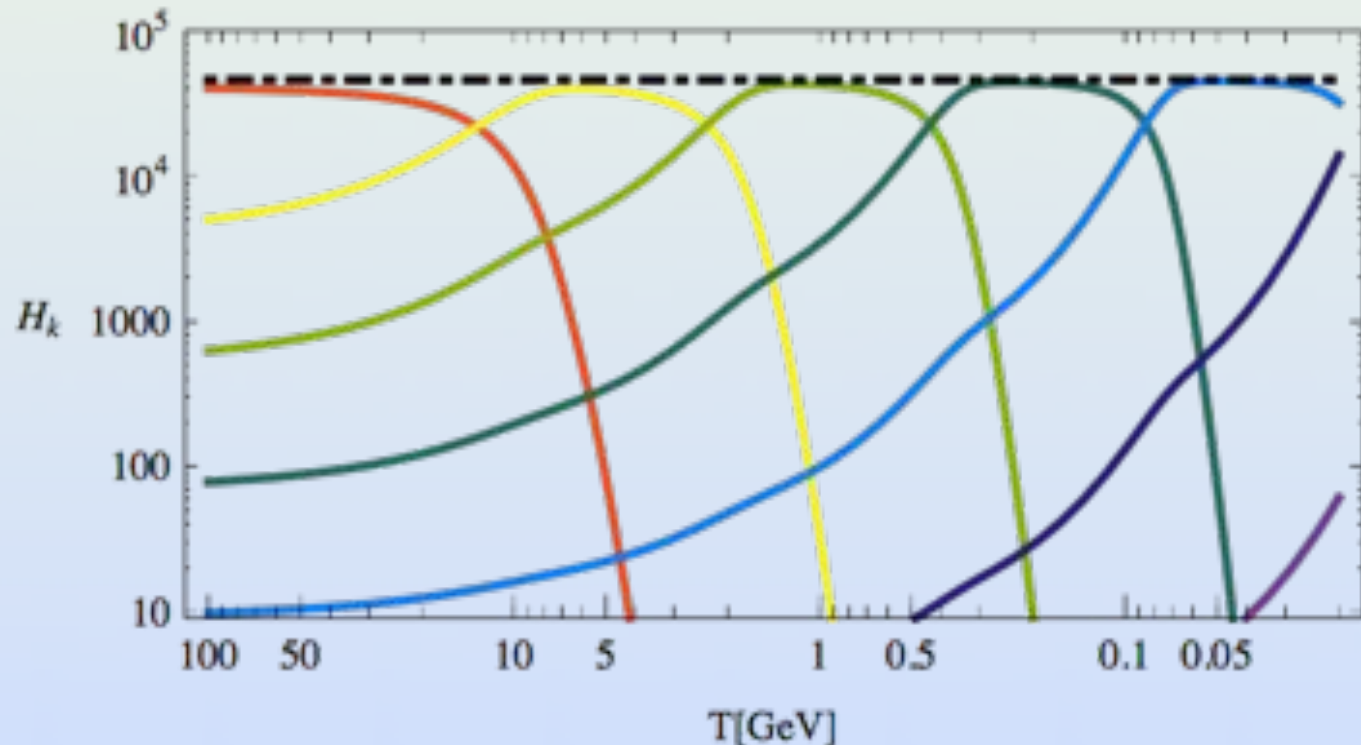
$$\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

[Vilenkin, Phys. Rev. D22, 3080 (1980)]

[Joyce & Shaposhnikov, astro-ph/9703005]

[Giovannini & Shaposhnikov, hep-ph/9710234]

[Boyarsky et al., arXiv:1109.3350]



# Hall effect in graphene

- Abnormal quantum Hall effect in graphene (magnetic field induced QH states)

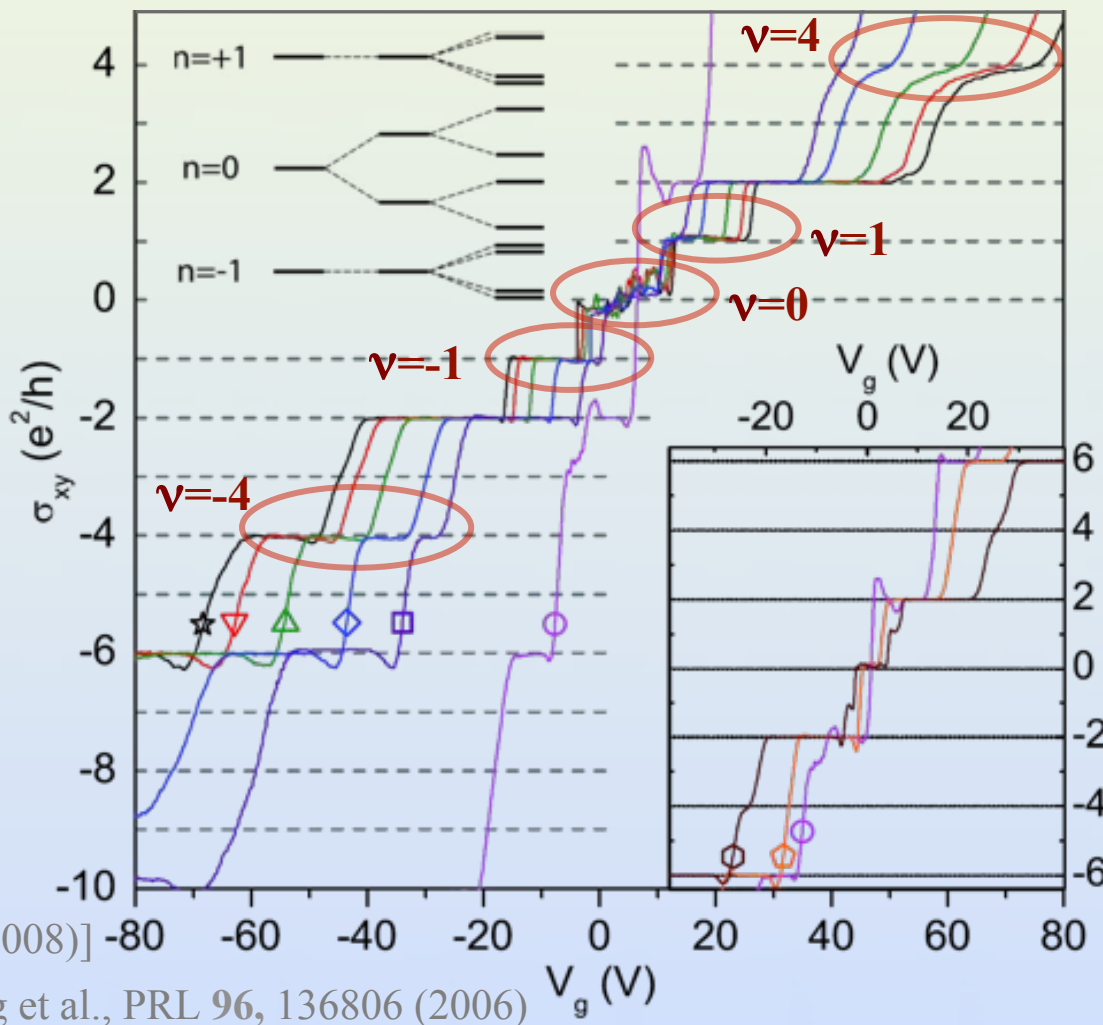
$$H_{\text{low-energy}} = v_F \int d^2r \bar{\Psi}_s \left( \gamma^1 \pi_x + \gamma^2 \pi_y \right) \Psi_s + H_{\text{int}}$$

- $B \rightarrow 0$ :

$$\sigma_{xy} = \nu \frac{e^2}{h} = 4 \left( n + \frac{1}{2} \right) \frac{e^2}{h}$$

- Broken symmetry  
QH states @ large B

$$\nu = 0, \pm 1, \pm 3, \pm 4$$



[Novoselov et al., Science **315**, 1379 (2007)]

[Abanin et al., Phys. Rev. Lett. **98**, 196806 (2007)]

[Checkelsky et al., Phys. Rev. Lett. **100**, 206801 (2008)]

[Xu Du et al., Nature **462**, 192 (2009)]

Zhang et al., PRL **96**, 136806 (2006)

# Dirac/Weyl semimetals

- Magnetoresistance in Dirac/Weyl semimetals (signs of chiral anomaly, dHvA oscillations)

“chiral shift”

$$H^{(W)} = \int d^3r \bar{\psi} \left[ -iv_F (\vec{\gamma} \cdot \vec{\mathbf{p}}) - (\vec{b} \cdot \vec{\gamma}) \gamma^5 - \mu \gamma^0 \right] \psi + H_{\text{int}}$$

- Anomalous contribution to conductivity:

$$\sigma_{33} = \sigma_{33}^{(LLL)} + \sigma_{33}^{(HLL)}, \quad \text{where} \quad \sigma_{33}^{(LLL)} = \frac{e^2 v_F |eB|}{4\pi^2 c \Gamma_0}$$

$$\sigma_{12} = \sigma_{12}^{(b=0)} + \sigma_{12,\text{anom}}, \quad \text{where} \quad \sigma_{12,\text{anom}} = \frac{e^2 b}{2\pi^2}$$

[Nielsen & Ninomiya, Phys. Lett. 130B, 390 (1983)]

[Burkov & Balents, Phys. Rev. Lett. 107, 127205 (2011)]

[V. Aji, Phys. Rev. B 85, 241101 (2012)]

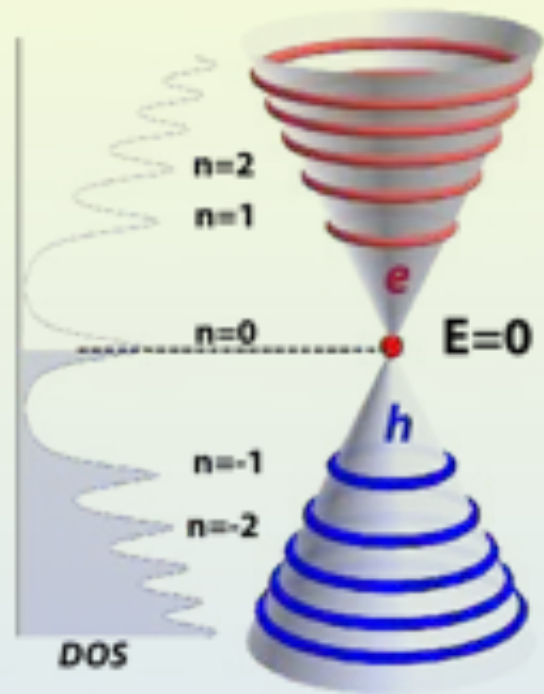
[Son & Spivak, Phys. Rev. B 88, 104412 (2013)]

[Gorbar, Miransky, Shovkovy, Phys. Rev. B 89, 085126 (2014)]

[Lu, Shen, Phys. Rev. B 92, 035203 (2015)]

[Klier, Gornyi, Mirlin, arXiv:1507.03481]

[Monteiro, Abanov, Kharzeev, arXiv:1507.05077]



# SOME TECHNICAL PROBLEM

# Schwinger propagator

- Proper-time representation in (2+1)-D

$$S(x, y) = \exp\left(-e \int_y^x A_\mu dz^\mu\right) \bar{S}(x - y)$$

$$\bar{S}(x) = \int \frac{d^3 k}{(2\pi)^3} e^{-ikx} \bar{S}(k)$$

Schwinger phase

where

$$\begin{aligned} \bar{S}(k) = & \int_0^\infty ds \exp\left(-is\left[m^2 - k_0^2 - \mathbf{k}^2 \frac{\tan(eBs)}{eBs}\right]\right) \\ & \times \left[ k_\nu \gamma^\nu + m - (k^2 \gamma^1 - k^1 \gamma^2) \tan(eBs) \right] \left[ 1 - \gamma^1 \gamma^2 \tan(eBs) \right] \end{aligned}$$

- Landau levels?

[Schwinger, Phys. Rev. 82, 664 (1951)]

- Schwinger propagator is equivalent to

$$\bar{S}(k) = i \exp\left(-\frac{\mathbf{k}^2}{|eB|}\right) \sum_{n=0}^{\infty} \frac{(-1)^n D_n(k)}{k_0^2 - m^2 - 2n|eB|}$$

where

$$D_n(k) = (k_0 \gamma^0 + m) \left[ P_+ L_n\left(\frac{2\mathbf{k}^2}{|eB|}\right) - P_- L_{n-1}\left(\frac{2\mathbf{k}^2}{|eB|}\right) \right] + 4(\mathbf{k} \cdot \boldsymbol{\gamma}) L_{n-1}^1\left(\frac{2\mathbf{k}^2}{|eB|}\right)$$

Free propagator  $\rightarrow$  full dynamical propagator?

[Chodos, Everding, Owen, Phys. Rev. D42, 2881 (1990)]

- Ritus method

$$\bar{S}(x, y) = \sum_{n=0}^{\infty} \int \frac{dp_0 dp_2}{(2\pi)^3} E_p(x) \frac{1}{(\bar{\mathbf{p}} \cdot \boldsymbol{\gamma}) + \Sigma(p)} \bar{E}_p(y)$$

where  $\bar{\mathbf{p}} = \left( p_0, 0, -\sqrt{2|eB|n} \right)$  and  $\bar{E}_p(x) = \gamma^0 E_p^+(x) \gamma^0$

$$(\pi_\nu \gamma^\nu) E_p(x) = E_p(x) (\bar{\mathbf{p}} \cdot \boldsymbol{\gamma})$$

- Shortcomings:
  - Schwinger phase is never fully factorized
  - Eigenstates  $E_p(x)$  are Dirac matrices
  - Obscure meaning of parameters in  $\Sigma(p)$



# A better representation?

- Full propagator

$$G(x, y) = i \langle x | \left[ k_0 \gamma^0 + \hat{F}^+ (\boldsymbol{\pi} \cdot \boldsymbol{\gamma}) + \hat{\Sigma}^+ \right]^{-1} | y \rangle$$

Here  $\hat{F}^\pm$  &  $\hat{\Sigma}^\pm$  are functions of  $\gamma^0, i\gamma^1\gamma^2,$  &  $(\boldsymbol{\pi} \cdot \boldsymbol{\gamma})^2$

3 mutually commuting operators

$$\hat{F}^\pm = f \pm \gamma^0 g \pm i\gamma^1\gamma^2 \tilde{g} + i\gamma^0\gamma^1\gamma^2 \tilde{f}$$

$$\hat{\Sigma}^\pm = m \pm \gamma^0 \mu \pm i\gamma^1\gamma^2 \tilde{\mu} + i\gamma^0\gamma^1\gamma^2 \Delta$$

Dirac mass

Haldane mass

where  $f, g, \tilde{g}, \tilde{f}, m, \mu, \tilde{\mu}, \Delta$  depend on  $(\boldsymbol{\pi} \cdot \boldsymbol{\gamma})^2$

[Gorbar, Gusynin, Miransky, Shovkovy, Phys. Scripta T146, 014018 (2012)]

- Eigenstates for orbital motion:

$$\pi^2 |N p\rangle = |eB|(2N + 1) |N p\rangle$$

E.g., in the Landau gauge  $\mathbf{A}=(0, Bx)$ ,

$$\langle \mathbf{r} | N p \rangle = C_N H_N \left( \frac{x}{l} + pl \right) \exp \left( -\frac{(x + pl^2)^2}{2l^2} + ipy \right)$$

They satisfy

$$\int d^2 \mathbf{r} \langle N p | \mathbf{r} \rangle \langle \mathbf{r} | N' p' \rangle = \delta_{NN'} \delta(p - p')$$

$$\sum_{N=0}^{\infty} \int_{-\infty}^{\infty} dp \langle \mathbf{r} | N p \rangle \langle N p | \mathbf{r}' \rangle = \delta(\mathbf{r} - \mathbf{r}')$$

[Gorbar, Gusynin, Miransky, Shovkovy, Phys. Scripta T146, 014018 (2012)]

# New LL representation

- Proper-time representation in (2+1)-D

$$G(x, y) = \exp\left(-e \int_y^x A_\mu dz^\mu\right) \bar{G}(x - y)$$

$$\bar{G}(t, \mathbf{r}) = \int \frac{dk_0}{2\pi} e^{-ik_0 t} \bar{G}(k_0, \mathbf{r})$$

Schwinger phase

where

$$\bar{G}(k_0, \mathbf{r}) = i \frac{e^{-\xi/2}}{2\pi l^2} \sum_{n,\sigma,s} \left( \frac{s(k_0 + \mu_{n,\sigma}) - m_{n,\sigma}}{(k_0 + \mu_{n,\sigma})^2 - E_{n,\sigma}^2} \left[ \delta_{-\sigma}^s L_n(\xi) + \delta_{\sigma}^s L_{n-1}(\xi) \right] + \frac{i(f_{n,\sigma} - sg_{n,\sigma})}{(k_0 + \mu_{n,\sigma})^2 - E_{n,\sigma}^2} \frac{(\boldsymbol{\gamma} \cdot \mathbf{r})}{l^2} \right) P_{s,\sigma}, \quad \xi = \mathbf{r}^2 / (2l^2)$$

[Gorbar, Gusynin, Miransky, Shovkovy, Phys. Scripta T146, 014018 (2012)]

- Landau-level energies Haldane mass

$$n = 0: \quad E_{0,\sigma} = \Delta_0 + \sigma m_0$$

Dirac mass

$$n \geq 1: \quad E_{n,\sigma} = \sqrt{2(f_{n,\sigma}^2 - g_{n,\sigma}^2)n|eB| + m_{n,\sigma}^2}$$

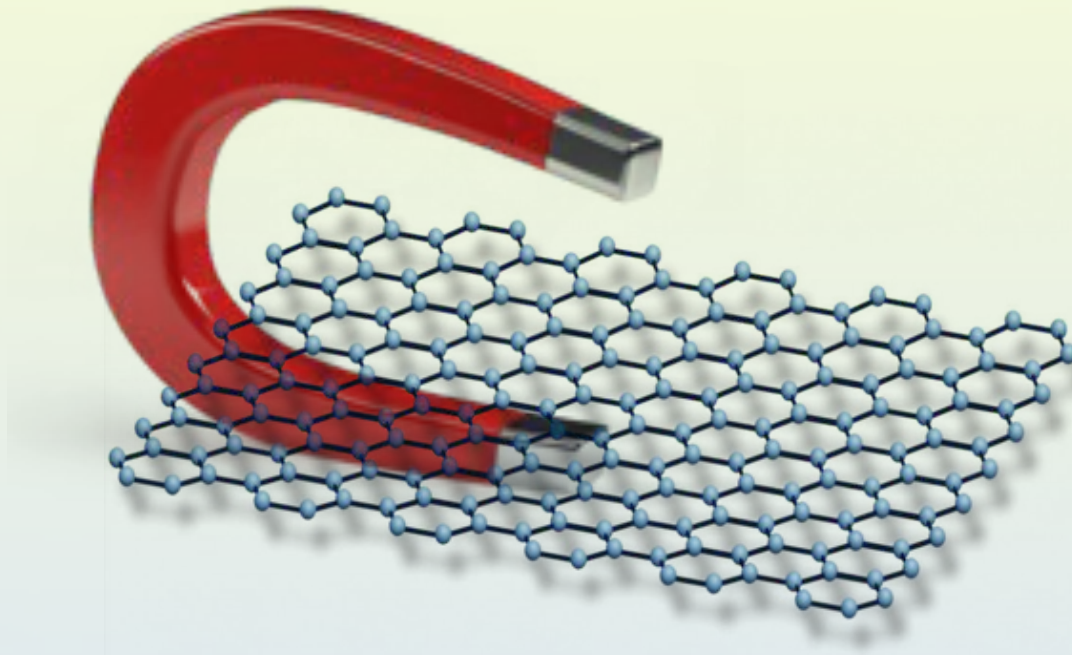
where  $m_{n,\sigma} = m_n + \sigma \Delta_n$ ,  $\mu_{n,\sigma} = \mu_n + \sigma \tilde{\mu}_n$ , etc.

By definition,  $n = N + (1 + s_{12})/2$  and

$$m\left((\boldsymbol{\pi} \cdot \boldsymbol{\gamma})^2\right) |N, s_{12}\rangle = m_n |N, s_{12}\rangle,$$

$$f\left((\boldsymbol{\pi} \cdot \boldsymbol{\gamma})^2\right) |N, s_{12}\rangle = f_n |N, s_{12}\rangle, \quad \text{etc.}$$

[Gorbar, Gusynin, Miransky, Shovkovy, Phys. Scripta T146, 014018 (2012)]



# **(2+1)-D GENERALIZED LANDAU-LEVEL REPRESENTATION**

# ASU Coulomb interaction in graphene

- Model Hamiltonian

U(4)

$$H = H_0 + H_{\text{Coulomb}} + \mu_B B \int d^2 r \Psi^\dagger \sigma_3 \Psi$$

SU(2)<sub>↑</sub> × SU(2)<sub>↓</sub>

- MC and QHF order parameters

$$m_{n,s} : \quad \bar{\Psi}_s \Psi_s \quad \text{Dirac mass (CDW)}$$

$$\Delta_{n,s} : \quad \bar{\Psi}_s (i\gamma^0 \gamma^1 \gamma^2) \Psi_s \quad \text{Haldane mass}$$

$$\tilde{\mu}_{n,s} : \quad \bar{\Psi}_s (i\gamma^1 \gamma^2) \Psi_s \quad \text{pseudospin density}$$

$$\mu_{n,3} : \quad \bar{\Psi}_\uparrow \gamma^0 \Psi_\uparrow - \bar{\Psi}_\downarrow \gamma^0 \Psi_\downarrow \quad \text{spin polarization}$$

- Note that  $\Psi_s^T = (\psi_{KAs}, \psi_{KBs}, \psi_{K'Bs}, \psi_{K'As})$

- No symmetry breaking parameters
- Renormalization of  $v_F$ :

$$n \geq 1: \quad E_n = v_F \sqrt{2n|eB|} \quad \rightarrow \quad E_n = f_n v_F \sqrt{2n|eB|}$$

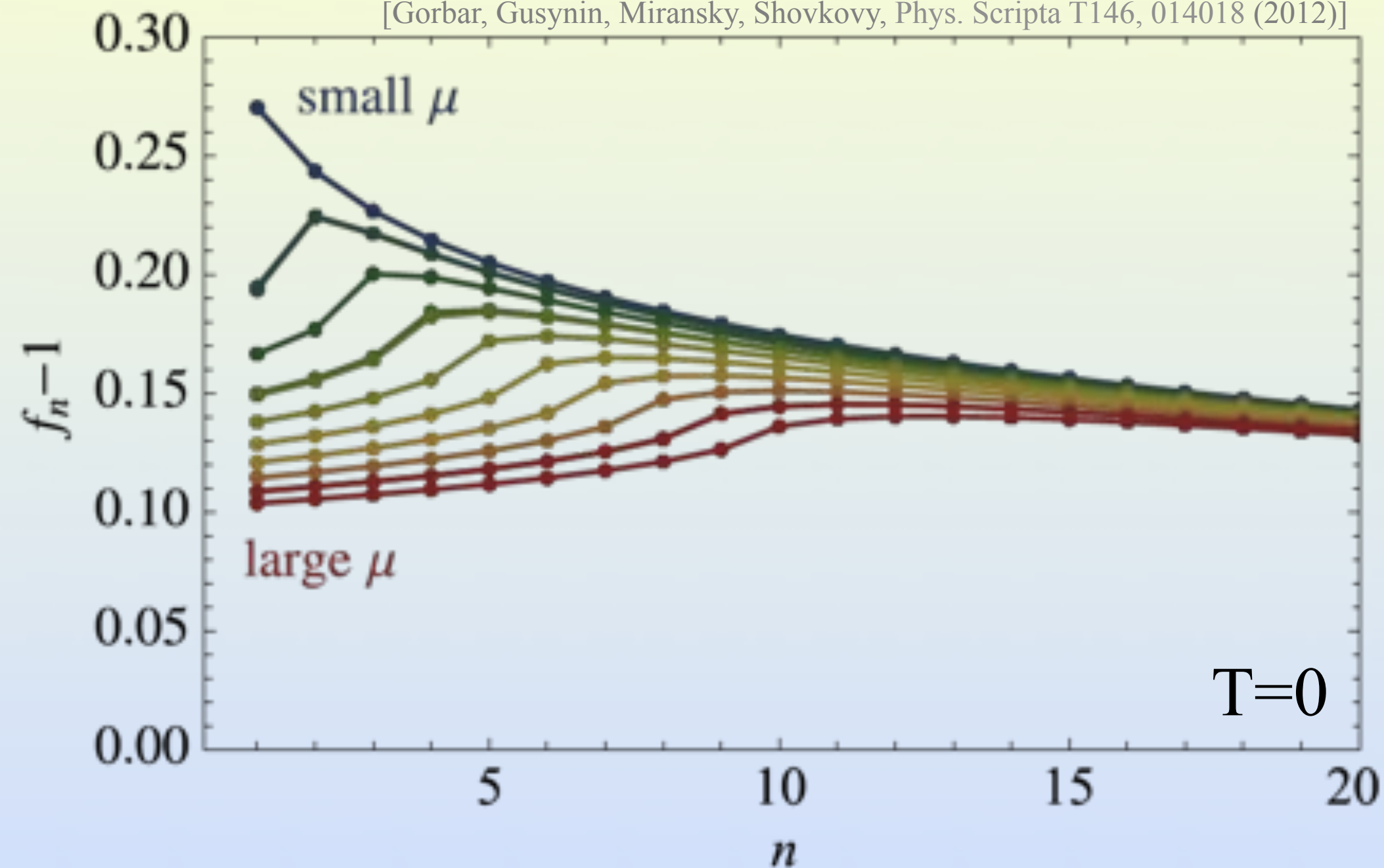
- Schwinger-Dyson equations for  $f_n$

$$f_n = 1 + \frac{\alpha}{2} \sum_{n=1}^{n_{\max}} \frac{K_{n'-1, n-1}^{(1)}}{n \sqrt{2n'}} \left[ 1 - n_F(E_{n'} - \mu) - n_F(E_{n'} + \mu) \right]$$

$$K_{m,n}^{(i)} = \int_0^\infty \frac{dk}{2\pi} \frac{kl L_{m,n}^{(i)}(kl)}{k + \Pi(0, k)}, \quad \text{where } l = 1 / \sqrt{|eB|}$$

Renormalization of  $\nu_F$ 

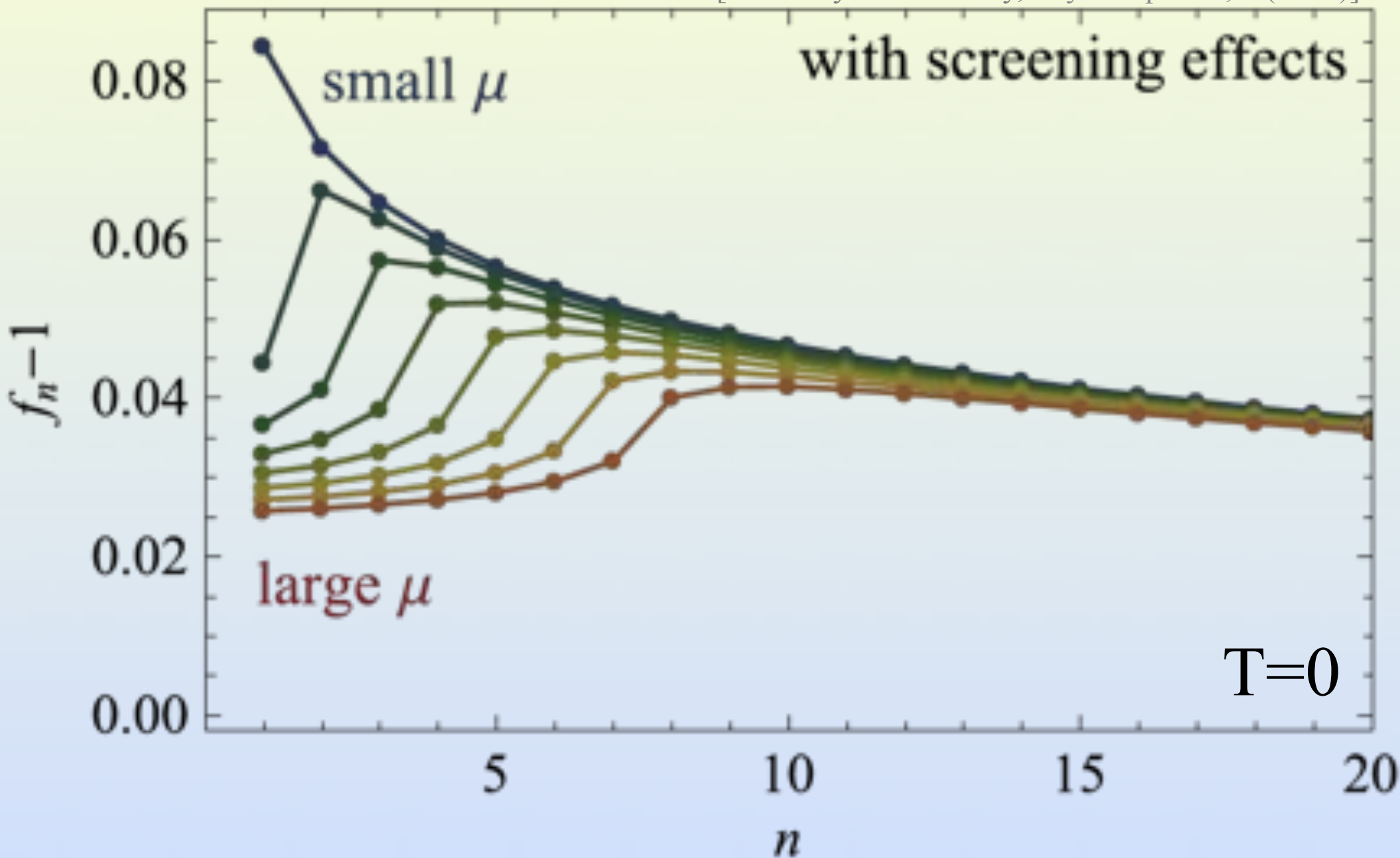
[Gorbar, Gusynin, Miransky, Shovkovy, Phys. Scripta T146, 014018 (2012)]



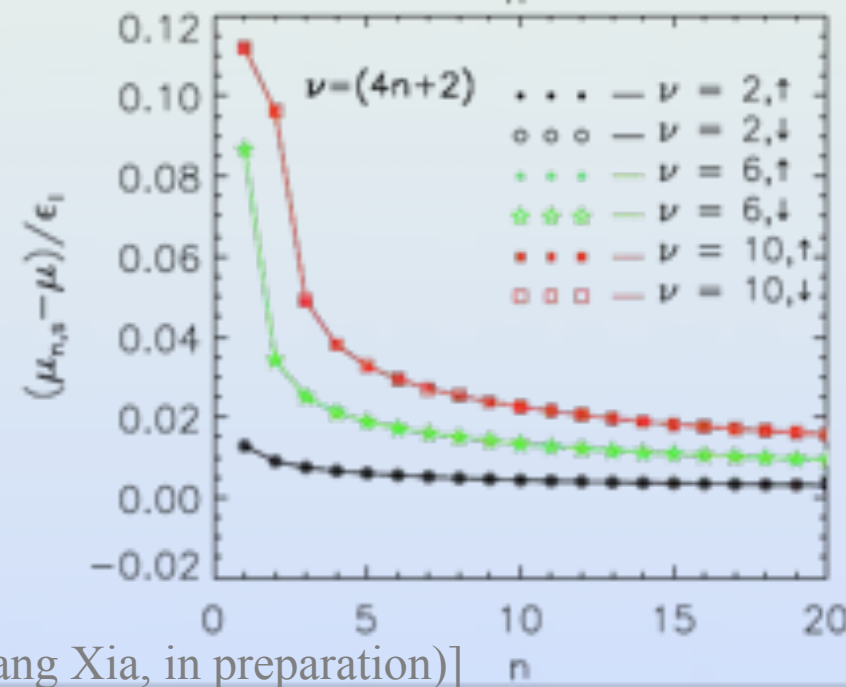
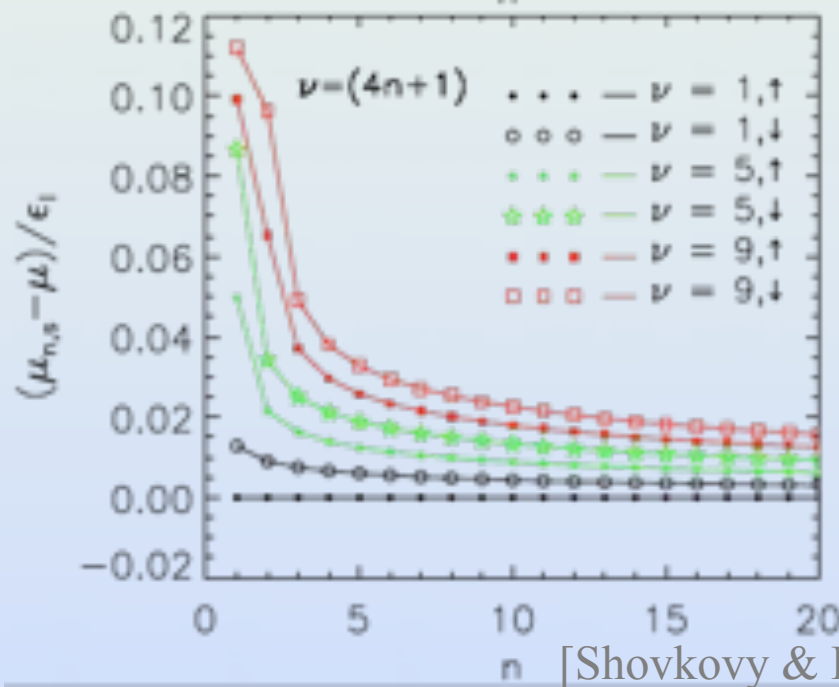
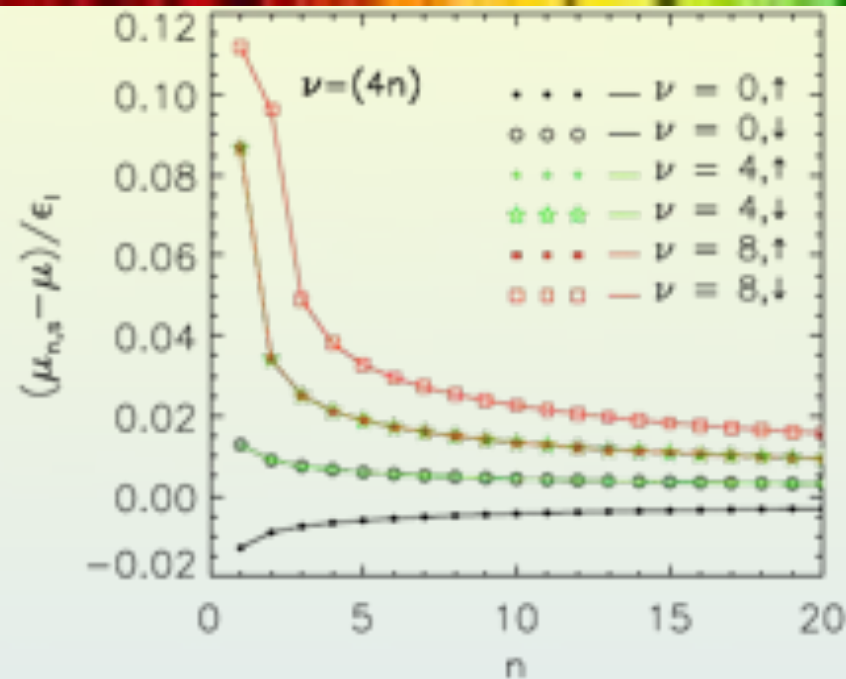
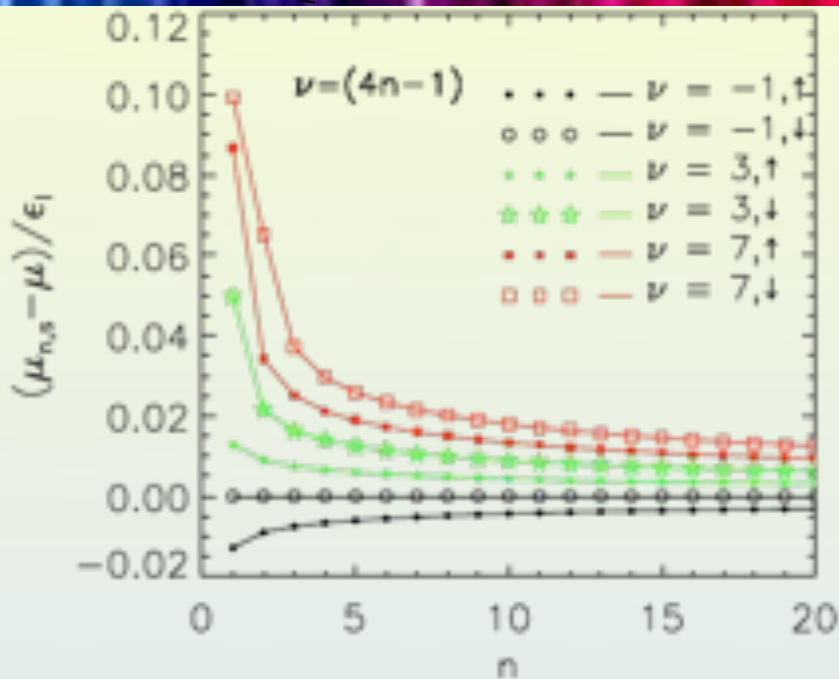


Renormalization of  $\nu_F$ 

[Miransky &amp; Shovkovy, Phys. Rep. 576, 1 (2015)]

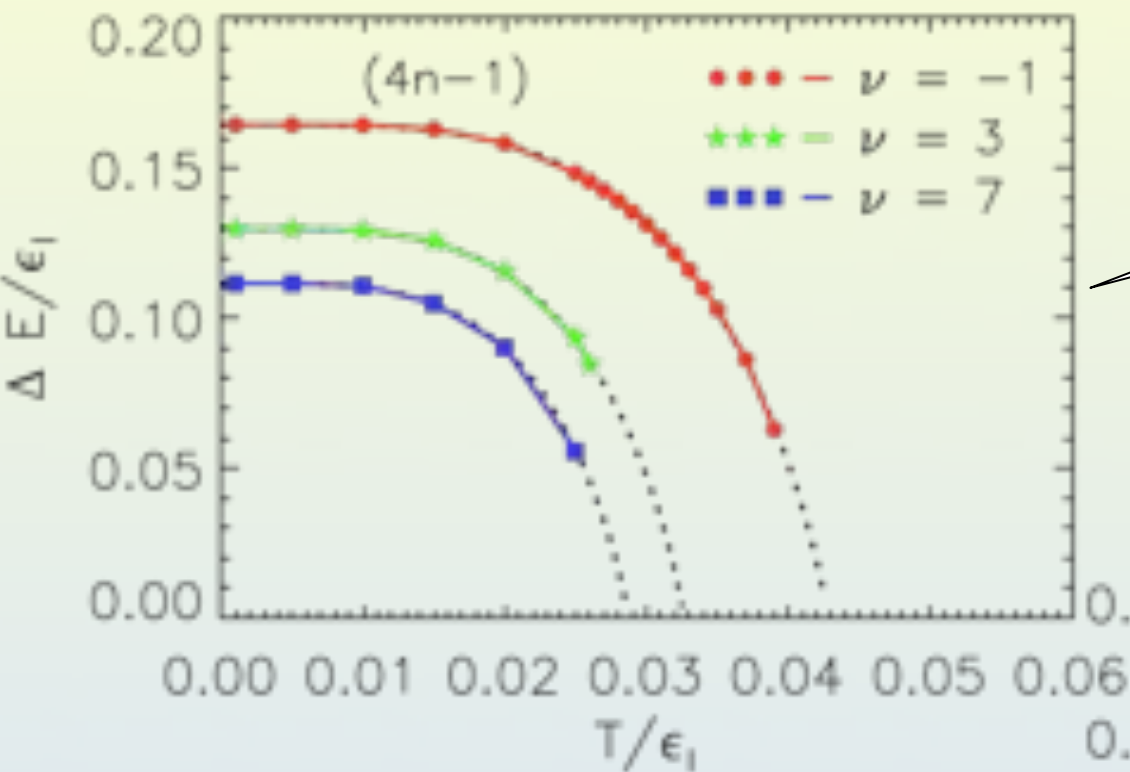


# QH states with different $\nu$



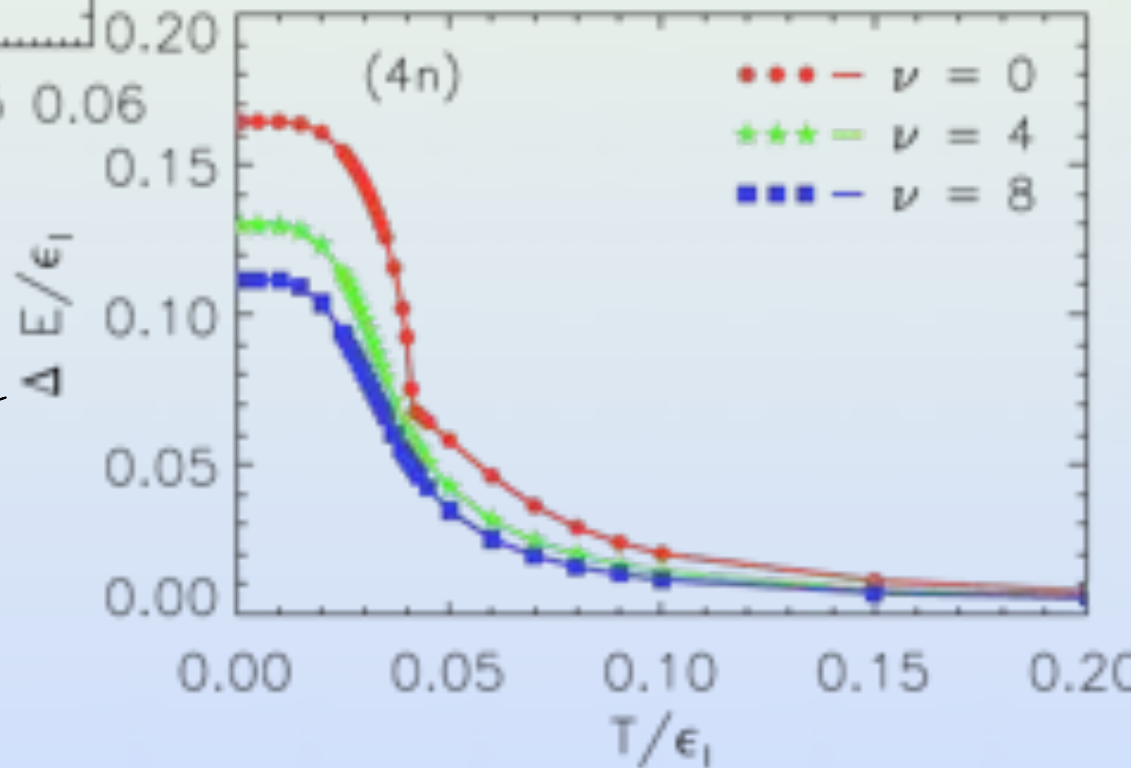
[Shovkovy & Lifang Xia, in preparation]

# Some results at $T \neq 0$



$U(1)_{\uparrow} \times SU(2)_{\downarrow}$

[Shovkovy & Lifang Xia, in preparation]



$SU(2)_{\uparrow} \times SU(2)_{\downarrow}$



# **(3+1)-D GENERALIZED LANDAU-LEVEL REPRESENTATION**

- Same as (inverse) propagator

$$\Sigma(x, y) = \exp\left(-e \int_y^x A_\mu dz^\mu\right) \bar{\Sigma}(x - y)$$

$$\bar{\Sigma}(x) = \int \frac{d^4 k}{(2\pi)^2} e^{-ikx} \bar{\Sigma}(k)$$

where

$$\bar{\Sigma} = 2e^{-\mathbf{k}_\perp^2 l^2} \sum_{n=0}^{\infty} (-1)^n \left( m_n + \gamma^3 \gamma^5 \Delta_n - \gamma^0 \gamma^5 \mu_{5,n} + \dots \right) \left[ P_+ L_n - P_- L_{n-1} \right] + \dots$$

Landau-level parameters  $m_n$ ,  $\Delta_n$ ,  $\mu_{5,n}$ , ... are given by LL-“projections”, e.g.,

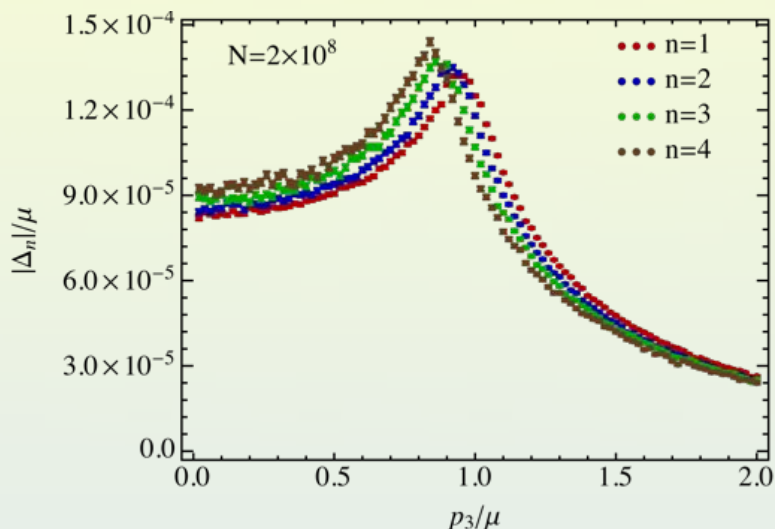
$$\Delta_n(k_0, k_3) = \frac{(-1)^n l^2}{8\pi} \int d^2 k_\perp e^{-k_\perp^2 l^2} [L_n + L_{n+1}] \text{Tr} [\gamma^0 \bar{\Sigma}(k)]$$

$$\mu_{5,n}(k_0, k_3) = \frac{(-1)^n l^2}{8\pi} \int d^2 k_\perp e^{-k_\perp^2 l^2} [L_n - L_{n+1}] \text{Tr} [\gamma^0 \gamma^5 \bar{\Sigma}(k)]$$

where, at 1-loop order,

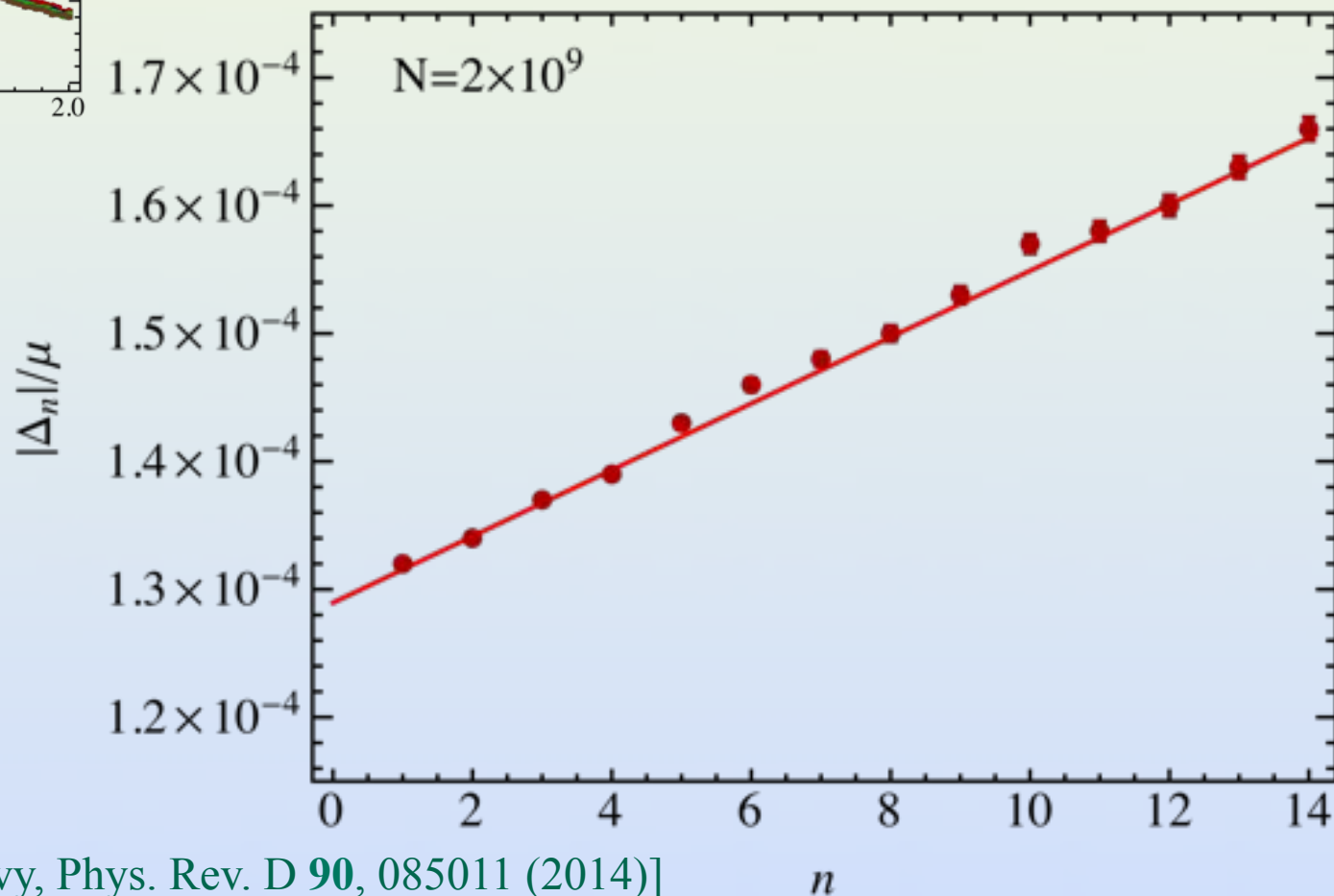
$$\bar{\Sigma}(p) = -4i \pi \alpha \int \frac{d^4 k}{(2\pi)^4} \gamma^\mu \bar{S}(k) \gamma^\nu D_{\mu\nu}(k-p)$$

[Gorbar, Miransky, Shovkovy, Wang, Phys. Rev. D **88**, 025043 (2013)]



Model fit:

$$\Delta_n = -\frac{\alpha |eB|}{\mu} \left( 0.53 + 0.32 \frac{|eB| n}{\mu^2} \right)$$

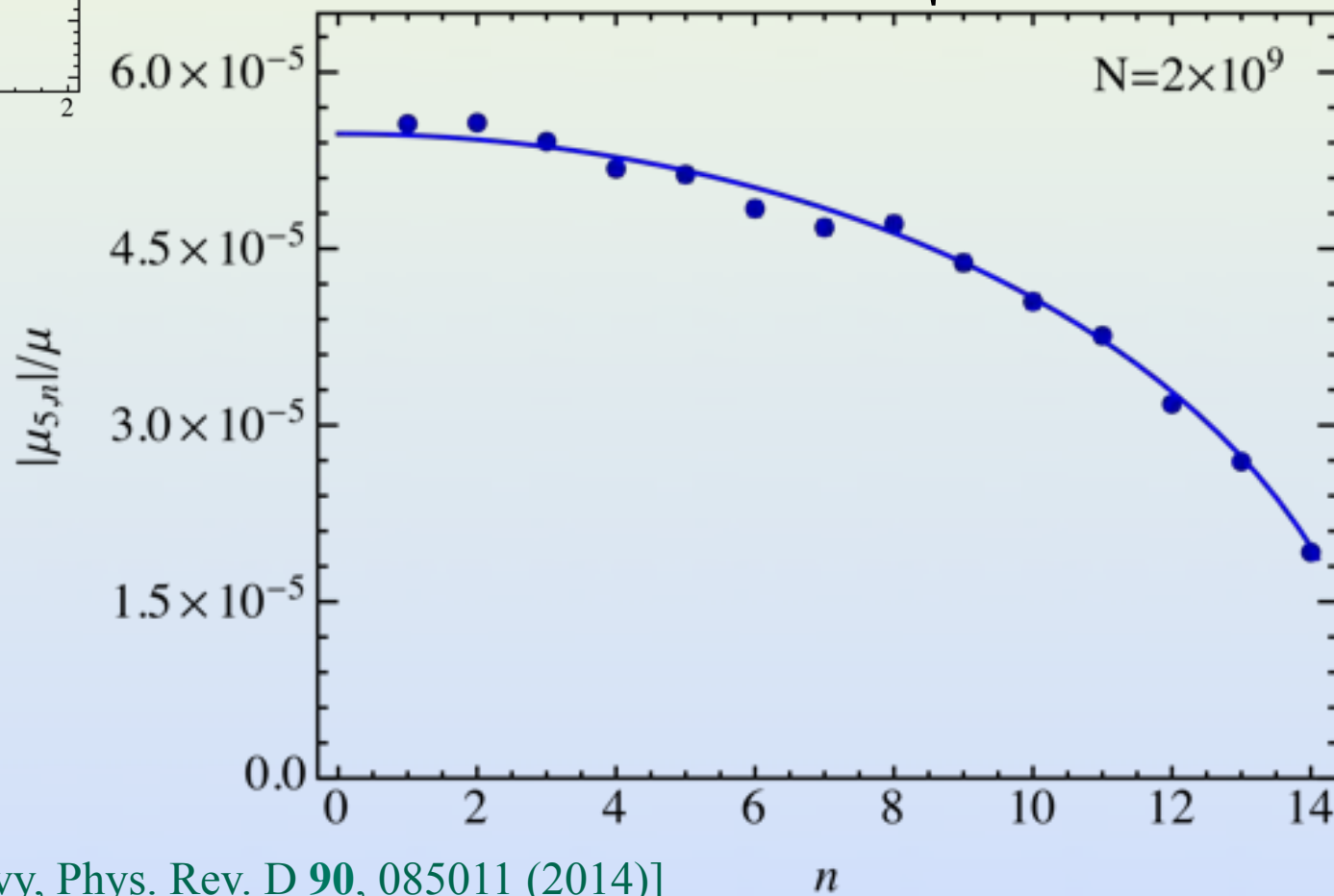
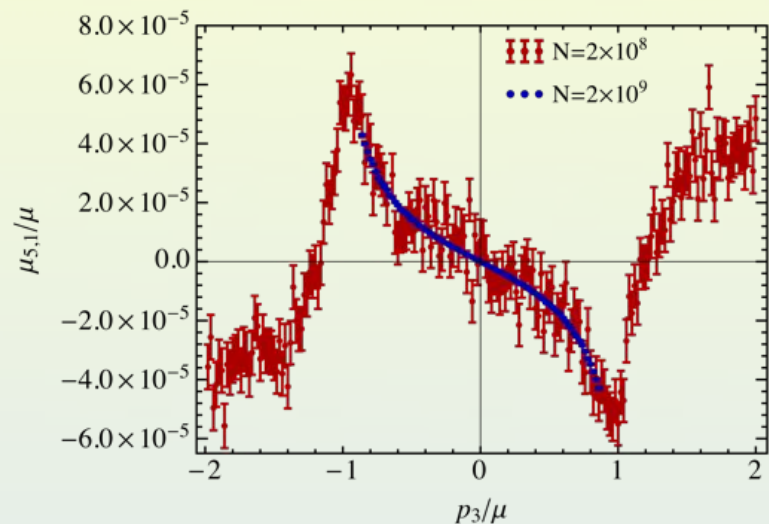


[Xia, Gorbar, Miransky, Shovkovy, Phys. Rev. D **90**, 085011 (2014)]

# $\mu_{5,n}$ in QED @ $B \neq 0$

Model fit:

$$\mu_{5,n} = -0.225 \frac{\alpha |eB|}{\mu} \sqrt{1 - \left( \frac{2n |eB|}{\mu^2} \right)^2}$$

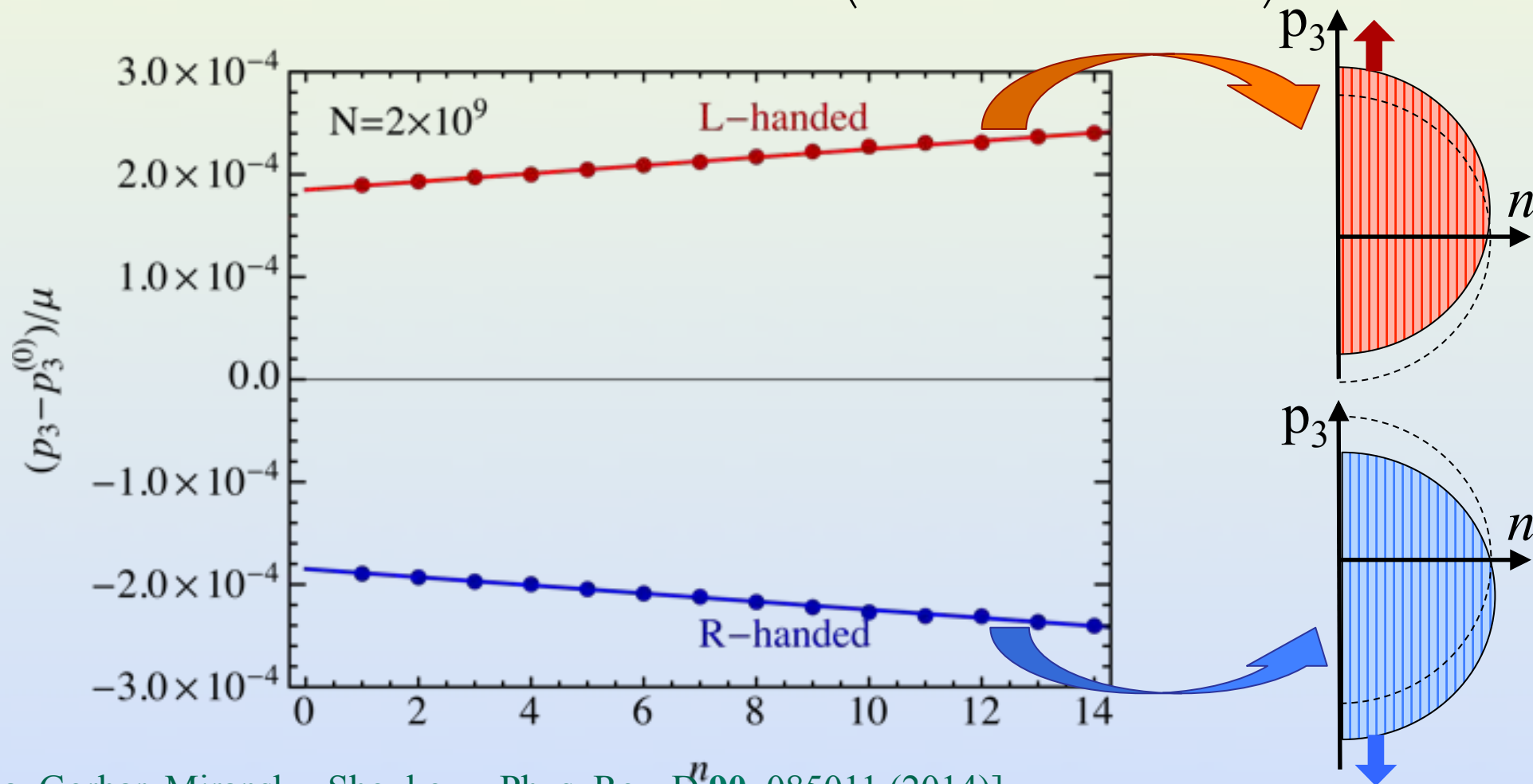


[Xia, Gorbar, Miransky, Shovkovy, Phys. Rev. D **90**, 085011 (2014)]



Model fit:

$$p_3 - p_3^{(0)} = \pm \frac{\alpha |eB|}{\mu} \left( 0.76 + 0.49 \frac{|eB| n}{\mu^2} \right)$$



[Xia, Gorbar, Miransky, Shovkovy, Phys. Rev. D **90**, 085011 (2014)]

# How large is the asymmetry?

In QED:

$$\frac{\alpha |eB|}{\mu} \approx 0.4 \left( \frac{B}{10^{18} \text{ G}} \right) \left( \frac{100 \text{ MeV}}{\mu} \right) \text{ MeV}/c$$

In QCD:

$$\frac{\alpha_s |eB|}{\mu} \approx 10 \left( \frac{B}{10^{18} \text{ G}} \right) \left( \frac{400 \text{ MeV}}{\mu} \right) \text{ MeV}/c$$

may have some consequences for compact stars...

- Many relativistic systems where nonzero  $B$  plays an essential role
- New generalized LL-representation is proposed
- Works in 2+1 as well as in 3+1 dimensions
- QH states in graphene can be described with highly flexible parametrization
- Chiral asymmetry in QED at  $B \neq 0$