



Anomalous chiral matter: from quark-gluon plasma to novel materials

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- **Early Universe, e.g.,**

[Boyarsky, Frohlich, Ruchayskiy, Phys.Rev.Lett. 108, 031301 (2012)]

- **Heavy-ion collisions, e.g.,**

[Kharzeev, Liao, Voloshin, Wang, Prog.Part.Nucl.Phys. 88, 1 (2016)]

- **Super-dense matter in compact stars, e.g.,**

[Yamamoto, Phys.Rev. D93, 065017 (2016)]

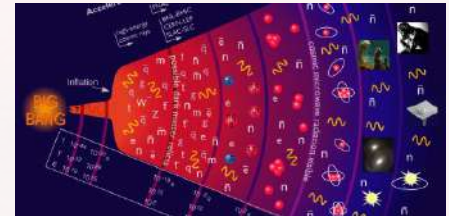
- **Ultra-relativistic jets from black holes**

- **Superfluid $^3\text{He-A}$, e.g.,**

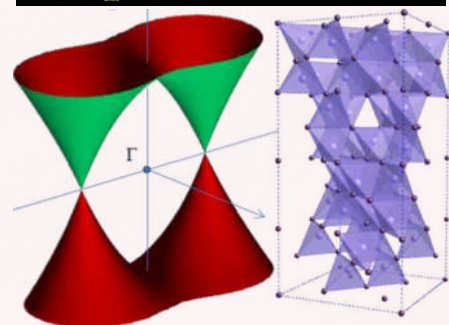
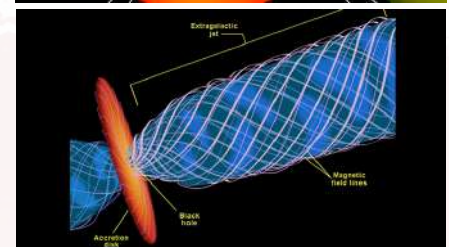
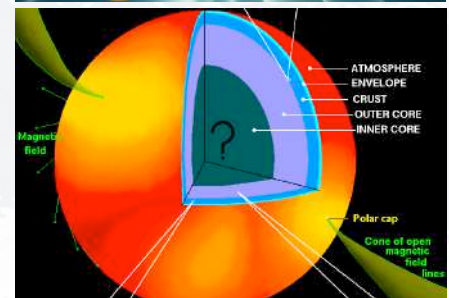
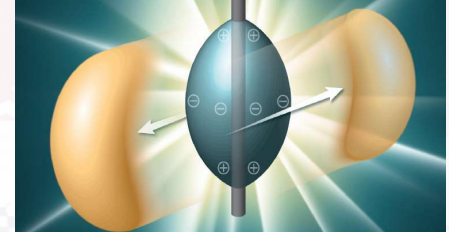
[Volovik, JETP Lett. 105, 34 (2017)]

- **Dirac/Weyl (semi-)metals, e.g.,**

[Li et. al. Nature Phys. 12, 550 (2016)]



Credit: Brookhaven National Laboratory



- *Massless* Dirac fermions:

$$\left(\gamma^0 p_0 - \vec{\gamma} \cdot \vec{p}\right) \Psi = 0 \quad \Rightarrow \quad \frac{\vec{\Sigma} \cdot \vec{p}}{|\vec{p}|} \Psi = \text{sign}(p_0) \gamma^5 \Psi$$

For particles ($p_0 > 0$): chirality = helicity

For antiparticles ($p_0 < 0$): chirality = - helicity

- Massive Dirac fermions in *ultrarelativistic* regime

– High temperature: $T \gg m$

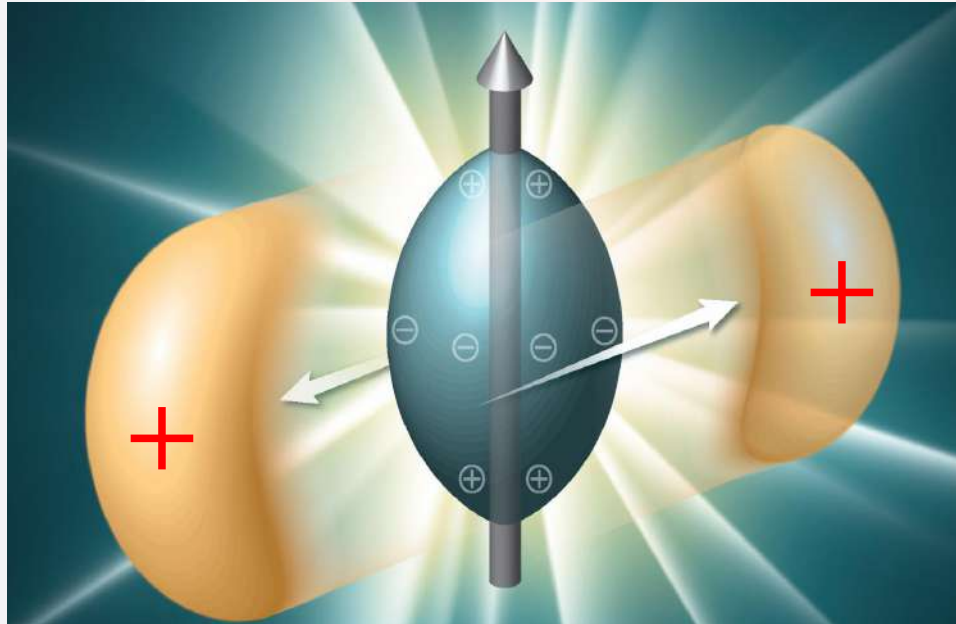
– High density: $\mu \gg m$

- Matter made of chiral fermions may allow $n_L \neq n_R$
- The chiral charge ($n_R - n_L$), unlike the electric charge ($n_R + n_L$), is **not** conserved

$$\frac{\partial(n_R + n_L)}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0$$

$$\frac{\partial(n_R - n_L)}{\partial t} + \vec{\nabla} \cdot \vec{j}_5 = \frac{e^2 \vec{E} \cdot \vec{B}}{2\pi^2 c}$$

- The chiral anomaly can have *macroscopic* implications in chiral matter



ANOMALOUS EFFECTS IN HEAVY-ION COLLISIONS

[Miransky & Shovkovy, Phys. Rep. **576**, 1 (2015)]

[Kharzeev, Liao, Voloshin, Wang, Prog. Part. Nucl. Phys. **88**, 1 (2016)]

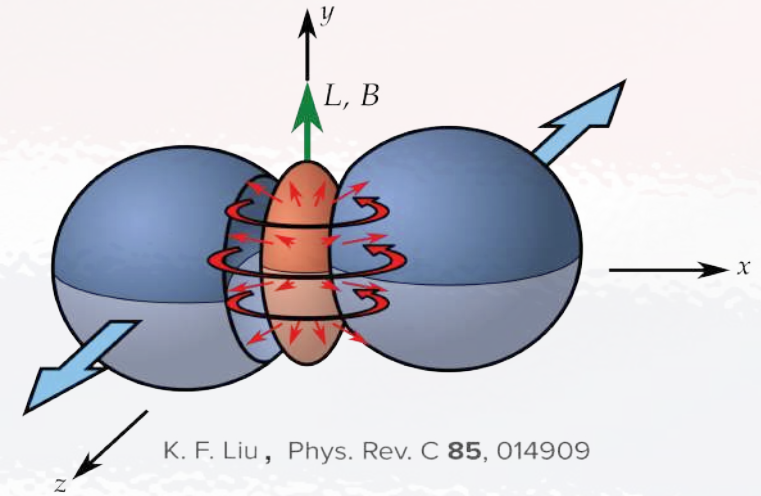
[Hattori & Huang, Nucl. Sci. Tech. **28**, 26 (2017)]

\vec{B} and $\vec{\omega}$ in little Bangs

- Rotating & magnetized QGP created at RHIC/LHC
- Electromagnetic fields (Lienard-Wiechert potentials)

$$e\mathbf{E}(t, \mathbf{r}) = \frac{e^2}{4\pi} \sum_n Z_n \frac{\mathbf{R}_n - R_n \mathbf{v}_n}{(R_n - \mathbf{R}_n \cdot \mathbf{v}_n)^3} (1 - v_n^2)$$

$$e\mathbf{B}(t, \mathbf{r}) = \frac{e^2}{4\pi} \sum_n Z_n \frac{\mathbf{v}_n \times \mathbf{R}_n}{(R_n - \mathbf{R}_n \cdot \mathbf{v}_n)^3} (1 - v_n^2)$$



[Rafelski & Müller, PRL, 36, 517 (1976)], [Kharzeev et al., arXiv:0711.0950], [Skokov et al., arXiv:0907.1396], [Voronyuk et al., arXiv:1103.4239], [Bzdak & Skokov, arXiv:1111.1949], [Deng & Huang, arXiv:1201.5108]

- Magnetic field estimate:

$$B \sim 10^{18} \text{ to } 10^{19} \text{ G } (\sim 100 \text{ MeV})$$

- Vorticity estimate:

$$\omega \sim 10^{21} \text{ s}^{-1} (\sim 10 \text{ MeV})$$

Effect of magnetic field

- Dirac equation with massless fermions

$$\left[i\gamma^0 \partial_0 - i\vec{\gamma} \cdot (\vec{\nabla} + ie\vec{A}) \right] \Psi = 0$$

- Energy spectrum

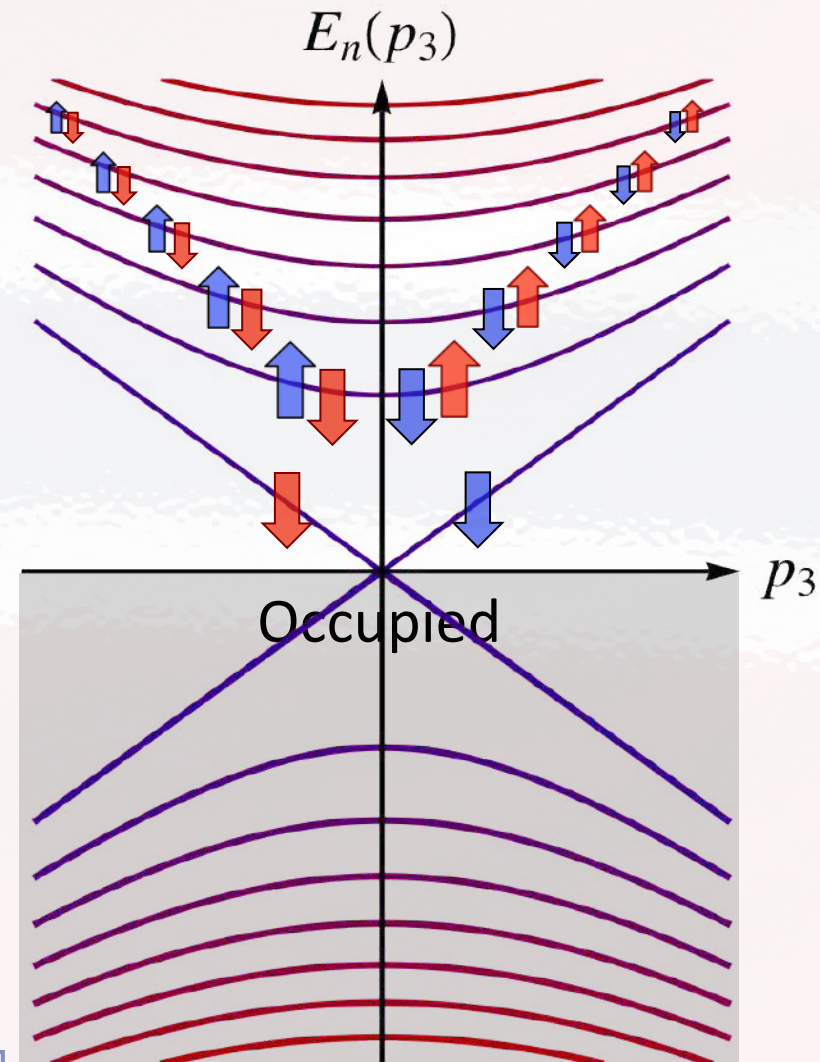
$$E_n^{(3+1)}(p_3) = \pm \sqrt{2n|eB| + p_3^2}$$

$$s = \pm \frac{1}{2} \quad (\text{spin})$$

where

$$n = s + k + \frac{1}{2}$$

$$k = 0, 1, 2, \dots \quad (\text{orbital})$$

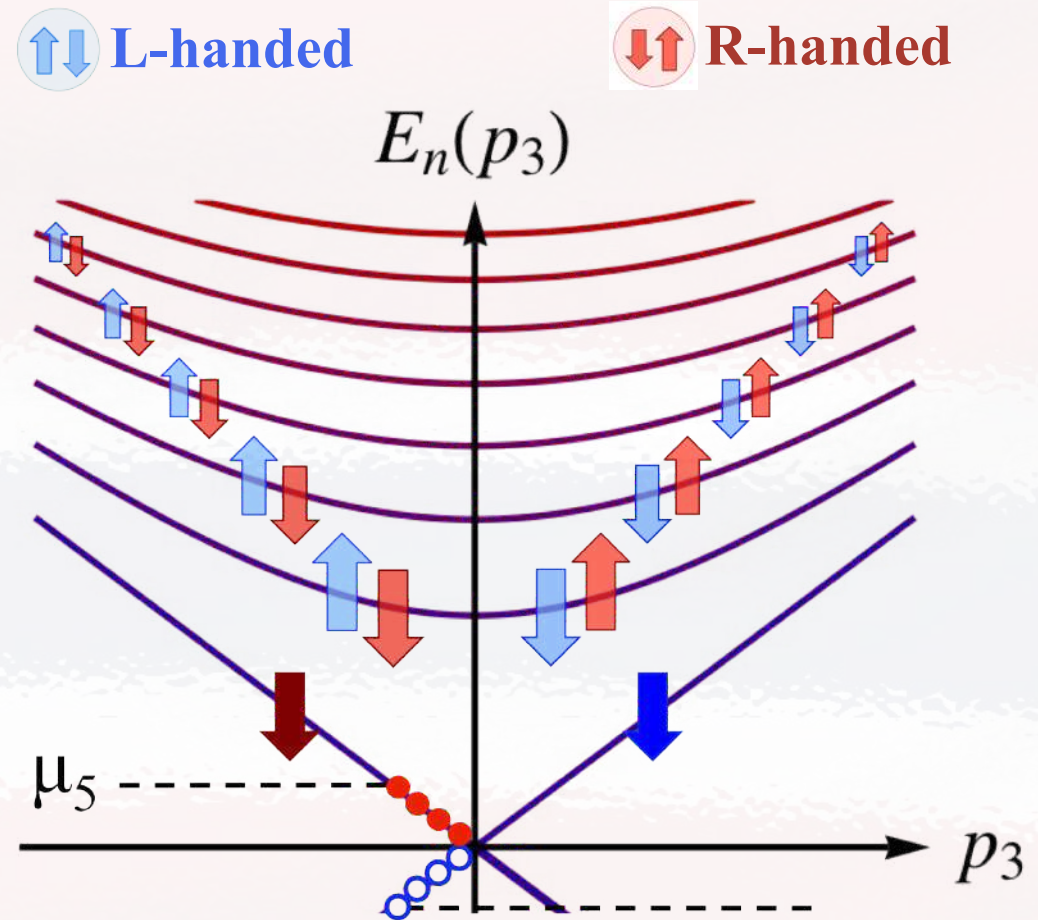


Chiral Magnetic Effect ($\mu_5 \neq 0$)

Topological fluctuations could induce *transient* state with a nonzero chiral charge ($\mu_5 \neq 0$)

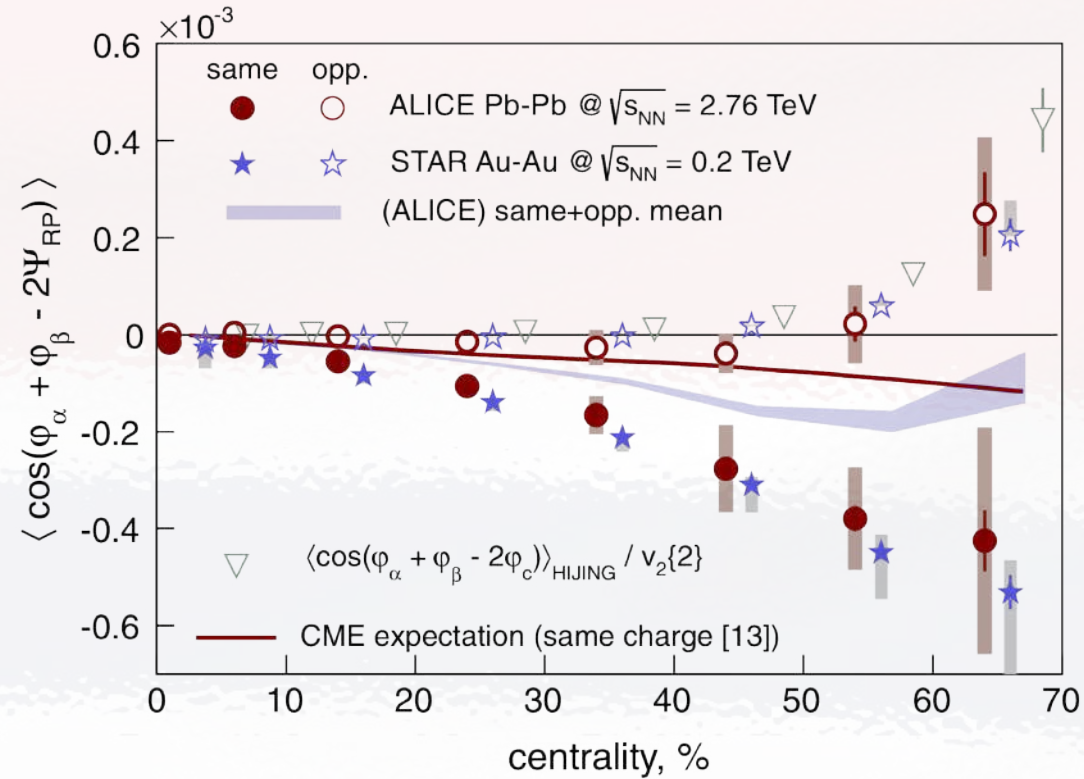
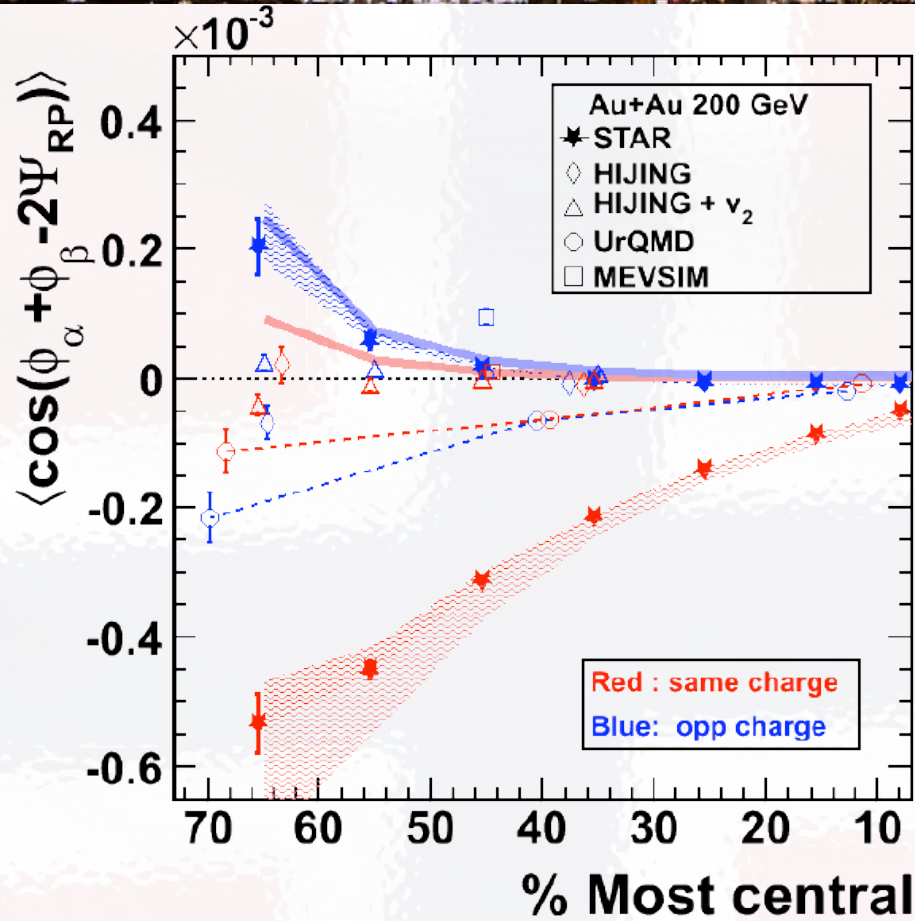
Spin polarized LLL ($s = \downarrow$ for particles of a *negative* charge):

- R-handed states $p_3 < 0$ give current in $+z$ direction
- L-handed holes $p_3 < 0$ give current in $+z$ direction too!



CME current:
$$\langle \vec{j} \rangle = \frac{e^2 \vec{B}}{2\pi^2} \mu_5$$

[Fukushima, Kharzeev, Warringa, Phys. Rev. D **78**, 074033 (2008)]



Correlations of same & opposite charge particles:

- [Abelev et al. (STAR), PRL **103**, 251601 (2009)]
- [Abelev et al. (STAR), PRC **81**, 054908 (2010)]
- [Abelev et al. (ALICE), PRL **110**, 012301 (2013)]
- [Adamczyk et al. (STAR), PRC **88**, 064911 (2013)]

$$\left\{ \begin{array}{l} \langle \cos(\varphi_\alpha^\pm + \varphi_\beta^\mp - 2\Psi_{RP}) \rangle \\ \langle \cos(\varphi_\alpha^\pm + \varphi_\beta^\pm - 2\Psi_{RP}) \rangle \end{array} \right.$$

Large background effects!

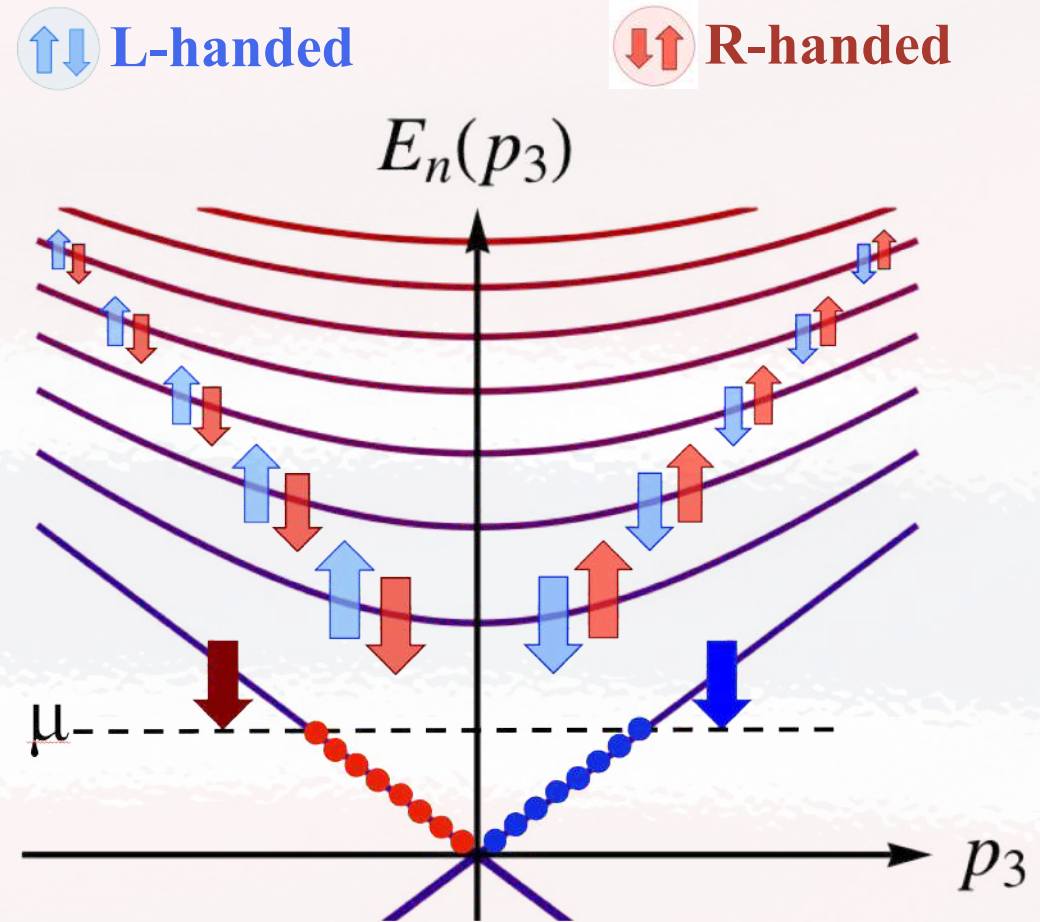
- [Belmont & Nagle, PRC **96**, 024901 (2017)]
- [ALICE Collaboration, Phys. Lett. B **777**, 151 (2018)]

Spin polarized LLL ($s=\downarrow$ for particles of a *negative* charge):

- R-handed states $p_3 < 0$
- L-handed states $p_3 > 0$

This gives rise to a nonzero axial current density (CSE):

$$\langle \vec{j}_5 \rangle = -\frac{e\vec{B}}{2\pi^2} \mu$$



[Vilenkin, Phys. Rev. D 22 (1980) 3067]

[Metlitski & Zhitnitsky, Phys. Rev. D 72, 045011 (2005)]

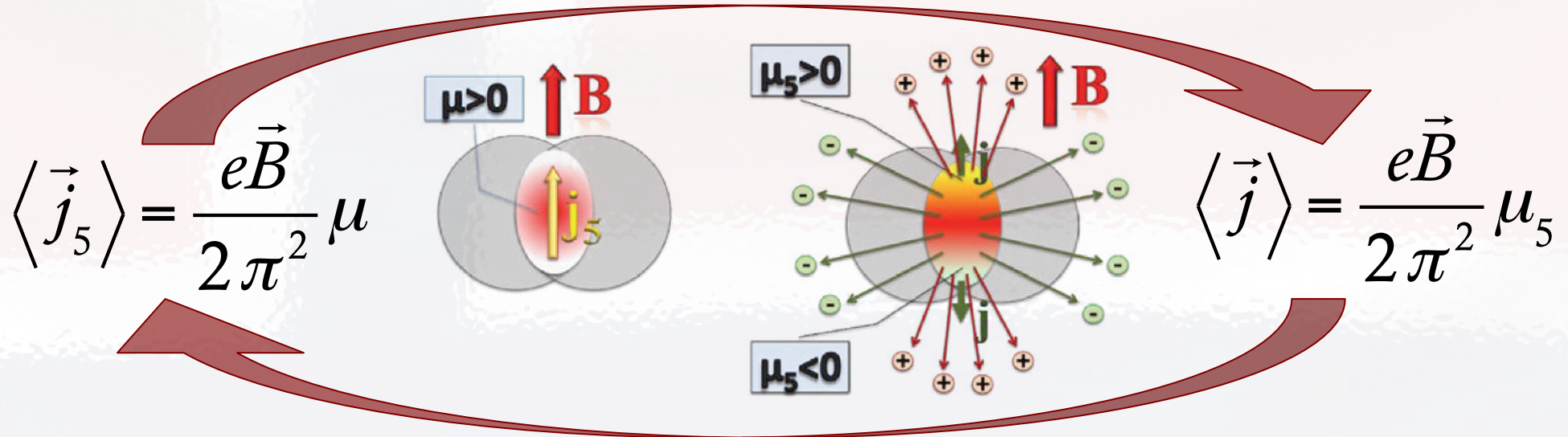
[Newman & Son, Phys. Rev. D 73 (2006) 045006]

Note that the latter can be also interpreted as a spin density:

$$\langle \vec{j}_5 \rangle = \langle \psi^\dagger \vec{\Sigma} \psi \rangle, \quad \text{where} \quad \Sigma^k = \frac{i}{2} \varepsilon^{klm} [\gamma_l, \gamma_m]$$

Chiral Magnetic Wave

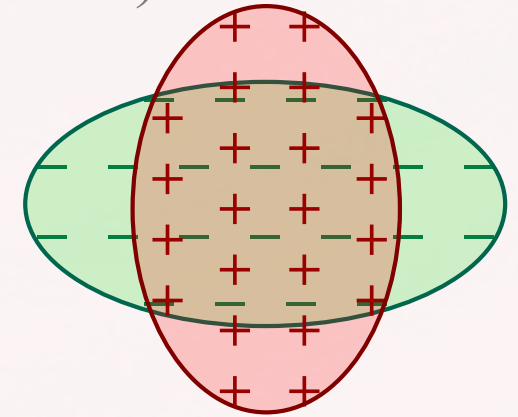
- Nonzero charge density @ $B \neq 0 \rightarrow$ CMW



[Gorbar, Miransky, Shovkovy, Phys. Rev. D **83**, 085003 (2011)]

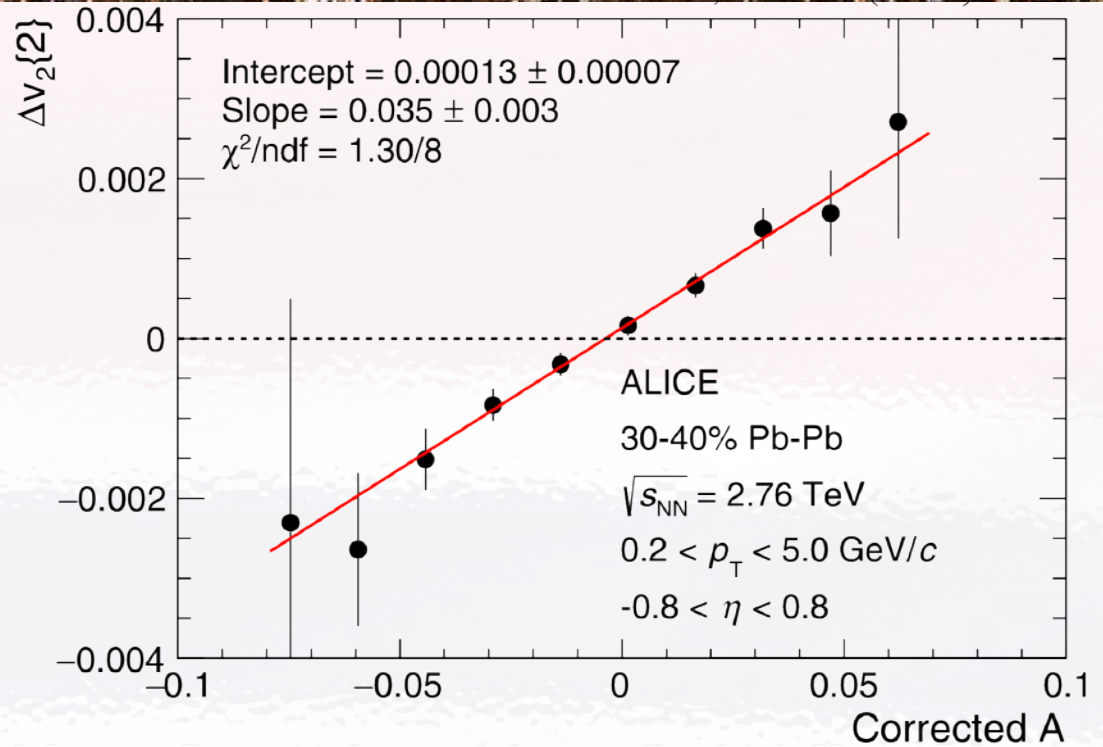
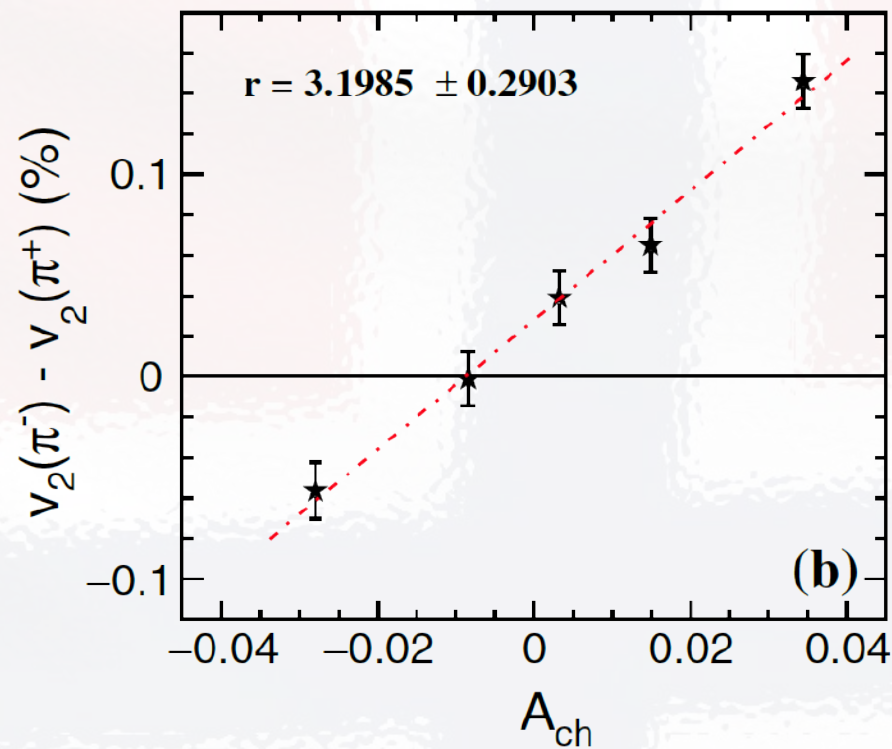
- Back-to-back electric currents, or quadrupole charge correlations (i.e., difference in elliptic flows of in π^+ and π^-)

$$\frac{dN_{\pm}}{d\phi} \approx \bar{N}_{\pm} [1 + 2v_2 \cos(2\phi) \mp A_{\pm} r \cos(2\phi)]$$



where A_{\pm} is the charge asymmetry

[Burnier, Kharzeev, Liao, Yee, Phys. Rev. Lett. **107** (2011) 052303]



[Ke (for STAR) J. Phys. Conf. Series **389**, 012035 (2012)]

[Adamczyk et al. (STAR), Phys. Rev. Lett. **114**, 252302 (2015)]

[Adam et al. (ALICE), Phys. Rev. C **93**, 044903 (2016)]

Higher harmonics of particle correlations are problematic...

Background effects may dominate over the signal!

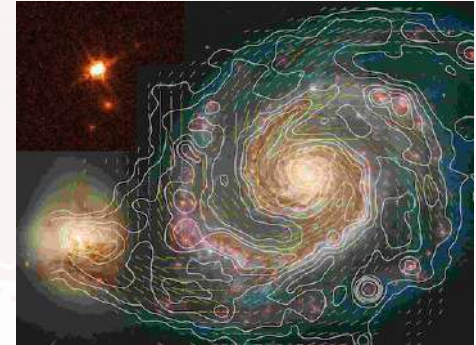
[CMS Collaboration, arXiv:1708.08901]

There are theoretical reasons that CMW is overdamped

[Rybalka, Gorbar, Shovkovy, arXiv:1807.07608]

Anomalous plasmas elsewhere?

- Inverse magnetic cascade may produce seeds of helical magnetic fields in the early Universe



[Vilenkin, Phys. Rev. D22, 3080 (1980)],
 [Joyce & Shaposhnikov, astro-ph/9703005],
 [Giovannini & Shaposhnikov, hep-ph/9710234]

- Eigenmodes of long wavelength and fixed helicity grow:

$$\frac{dB_k}{dt} \propto \frac{ck}{\sigma} (4\pi C_5 \mu_5 - ck) B_k$$

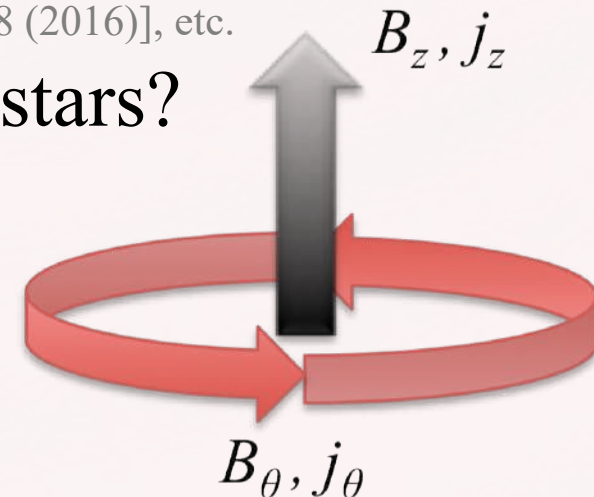
[Boyarsky et al., PRL **108**, 031301 (2012)], [Tashiro et al., PRD **86**, 105033 (2012)],
 [Manuel et al., PRD **92**, 074018 (2015)], [Hirono et al., PRD **92**, 125031 (2015)],
 [Buividovich et al., PRD **94**, 025009 (2016)], [Gorbar et al., PRD **94**, 103528 (2016)], etc.

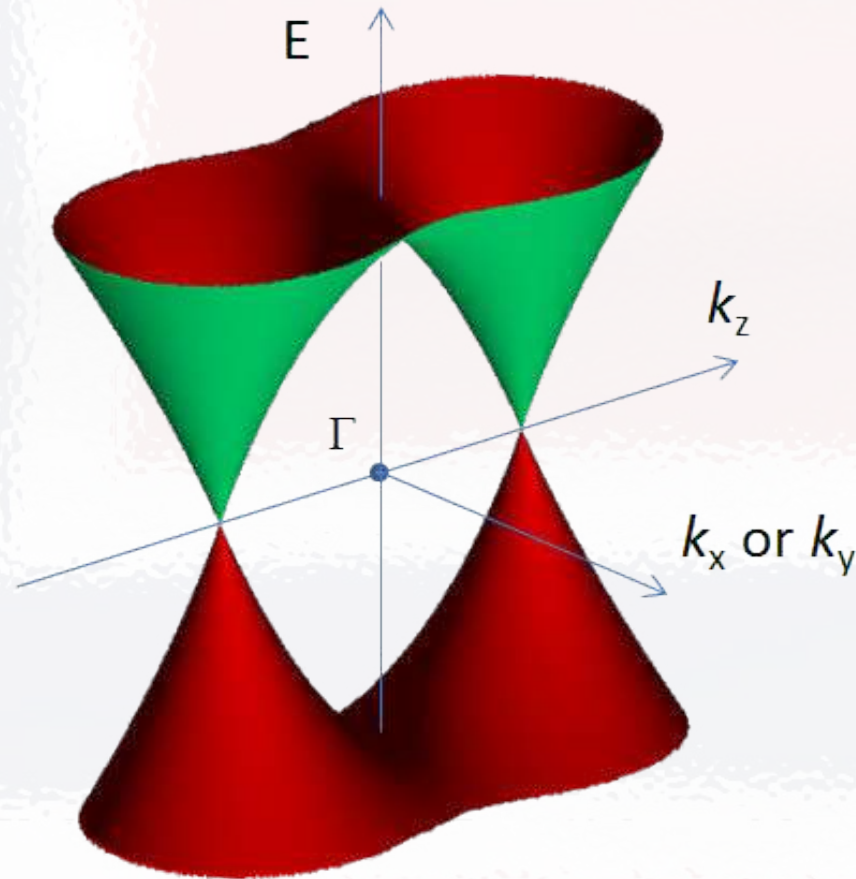
- Strong helical magnetic field in compact stars?

[Ohnishi, Yamamoto, arXiv:1402.4760]
 [Yamamoto, Phys. Rev. D **93**, 065017 (2016)]
 [Dvornikov, J. Exp. Theor. Phys. **123** 967 (2016)]

- Perhaps, not...

[Grabowska, Kaplan, Reddy, Phys. Rev. D **91**, 085035 (2015)]





Credits: Borisenko et al., Phys. Rev. Lett. 113, 027603 (2014)

DIRAC & WEYL MATERIALS

[Liu et al., Science **343**, 864 (2014)], [Neupane et al., Nature Commun. **5**, 3786 (2014)]
[Borisenko et al., Phys. Rev. Lett. **113**, 027603 (2014)], [Li et al., Nature Physics **12**, 550 (2016)]
[S.-Y. Xu et al., Science **349**, 613 (2015)], [B. Q. Lv et al., Phys. Rev. X **5**, 031013 (2015)]
[S.-Y. Xu et al., Nature Physics **11**, 748 (2015)], [S.-Y. Xu et al., Science Adv. **1**, 1501092 (2015)]
[F. Y. Bruno et al., Phys. Rev. B **94**, 121112 (2016)]

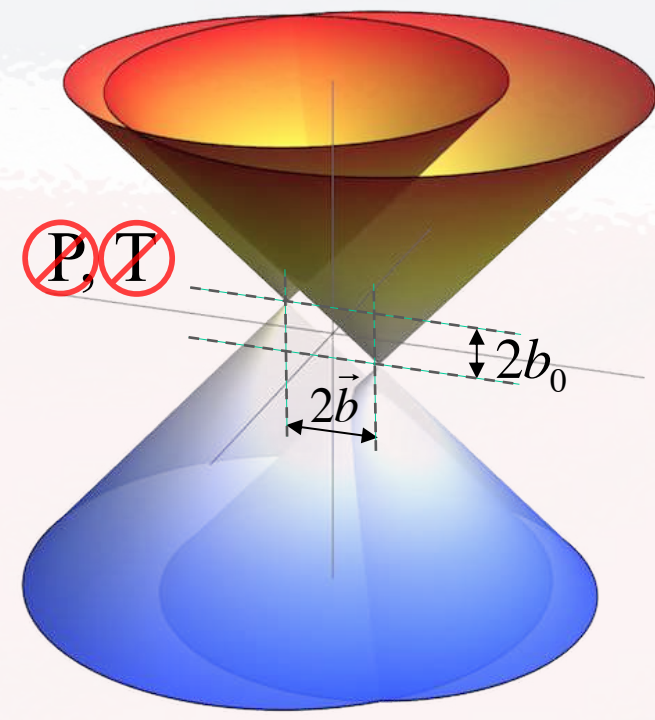
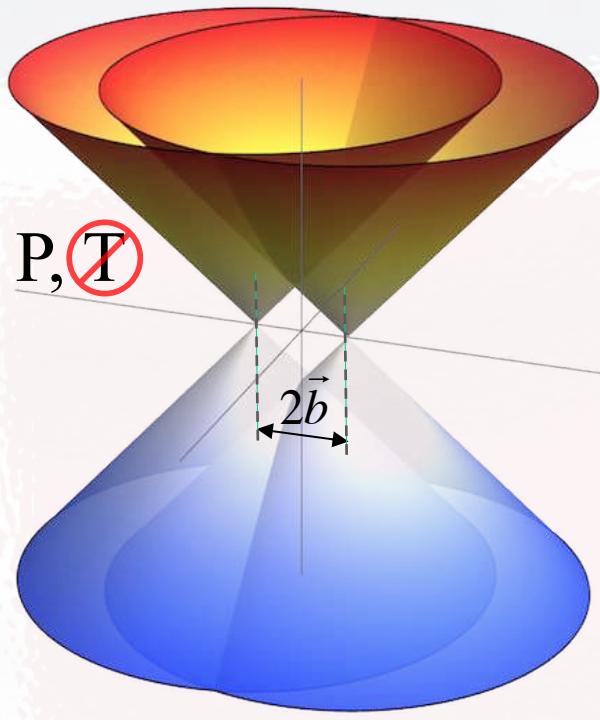
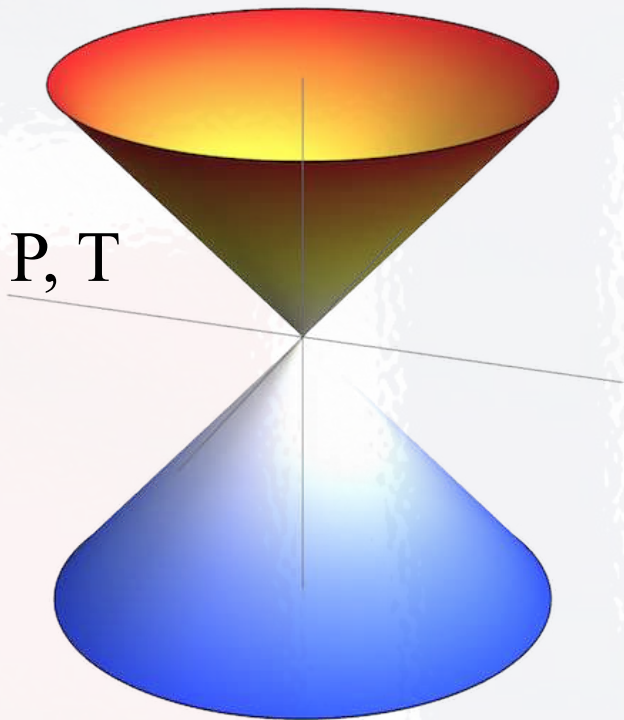
Dirac vs. Weyl materials

- Low-energy Hamiltonian of a Dirac/Weyl material

$$H = \int d^3\mathbf{r} \bar{\psi} \left[-iv_F (\vec{\gamma} \cdot \vec{\mathbf{p}}) - \underbrace{(\vec{b} \cdot \vec{\gamma})}_{\text{T}} \gamma^5 + \underbrace{b_0 \gamma^0 \gamma^5}_{\text{P}} \right] \psi$$

Dirac (e.g., Na₃Bi, Cd₃As₂, ZrTe₅)

Weyl (e.g., TaAs, NbAs, TaP, NbP, WTe₂)



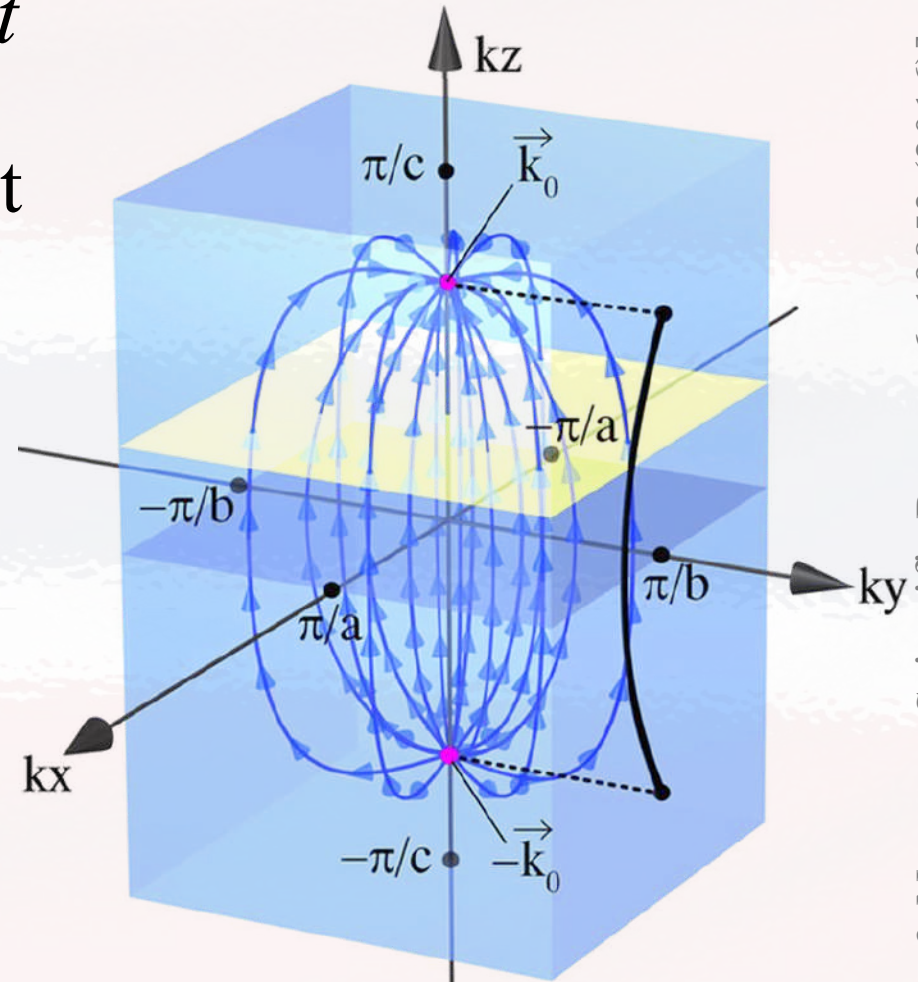
Weyl fermions on a lattice

- In solid state physics, the momentum space (Brillouin zone) is *compact*
- *Berry curvature* is nonzero at Weyl nodes

$$\vec{\Omega}_k = \lambda \frac{\vec{k}}{2k^3}$$

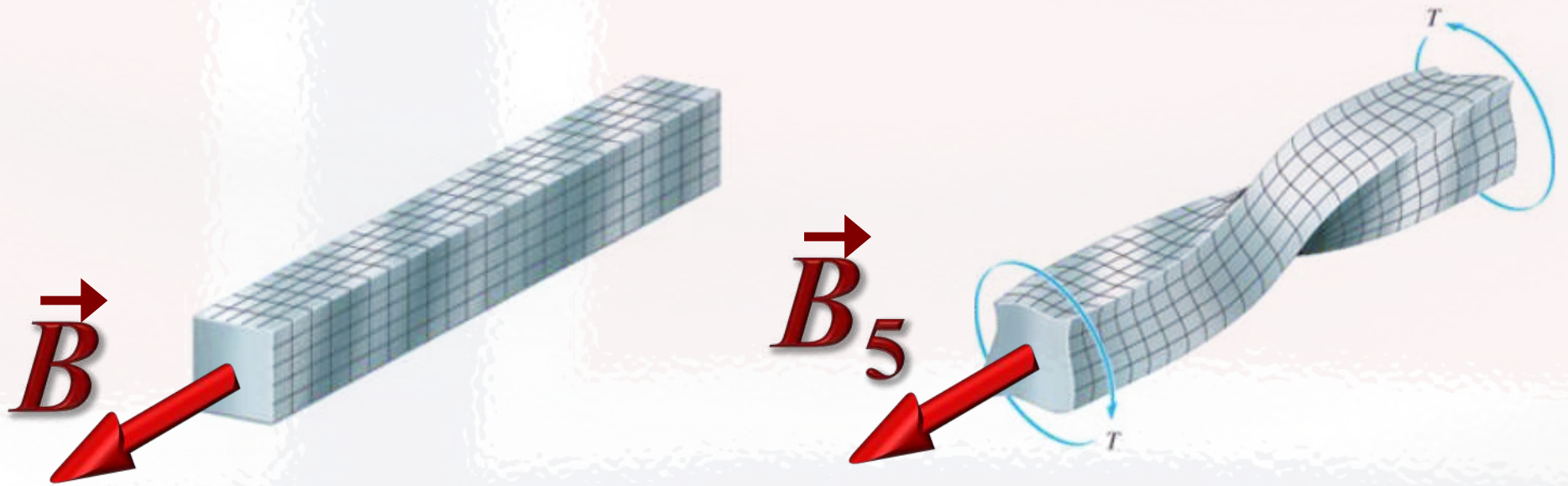
- Weyl nodes carry nonzero *topological charges*

$$C = \frac{1}{2\pi} \oint \vec{\Omega}_k \cdot d\vec{S} = \lambda = \pm 1$$



- Weyl fermions come in pairs of opposite chirality

[Nielsen & Ninomiya, Nucl. Phys. B **193**, 173 (1981); B **185**, 20 (1981)]



PSEUDO-ELECTROMAGNETIC FIELDS

[Zubkov, Annals Phys. **360**, 655 (2015)]

[Cortijo, Ferreira, Landsteiner, Vozmediano. Phys. Rev. Lett. **115**, 177202 (2015)]

[Grushin, Venderbos, Vishwanath, Ilan, Phys. Rev. X **6**, 041046 (2016)]

[Cortijo, Kharzeev, Landsteiner, Vozmediano, Phys. Rev. B **94**, 241405 (2016)]

[Pikulin, Chen, Franz, Phys. Rev. X **6**, 041021 (2016)]

- Strains in the low-energy effective Weyl Hamiltonian

$$H = \int d^3 \mathbf{r} \bar{\psi} \left[-i v_F (\vec{\gamma} \cdot \vec{\mathbf{p}}) - (\vec{b} + \vec{A}_5) \cdot \vec{\gamma} \gamma^5 + (b_0 + A_{5,0}) \gamma^0 \gamma^5 \right] \psi$$

where the chiral gauge fields are

$$A_{5,0} \propto b_0 \left| \vec{b} \right| \partial_{||} u_{||}$$

[Zubkov, Annals Phys. **360**, 655 (2015)]

[Cortijo, Ferreira, Landsteiner, Vozmediano. PRL **115**, 177202 (2015)]

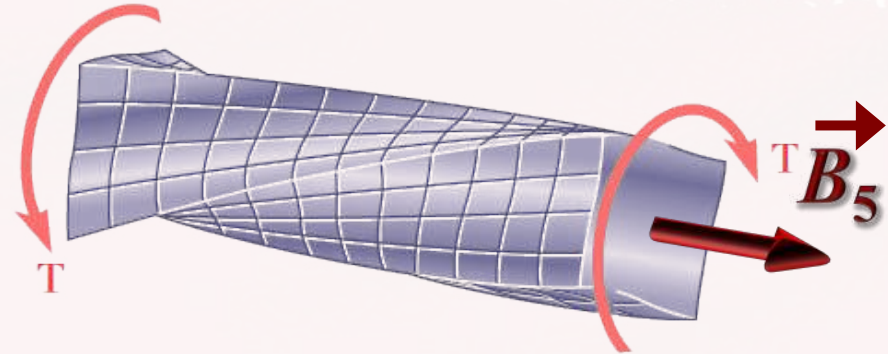
[Pikulin, Chen, Franz, PRX **6**, 041021 (2016)]

[Grushin, Venderbos, Vishwanath, Ilan, PRX **6**, 041046 (2016)]

$$A_{5,\perp} \propto \left| \vec{b} \right| \partial_{||} u_{\perp}$$

[Cortijo, Kharzeev, Landsteiner, Vozmediano, PRB **94**, 241405 (2016)]

$$A_{5,||} \propto \alpha \left| \vec{b} \right|^2 \partial_{||} u_{||} + \beta \sum_i \partial_i u_i$$



leading to the pseudo-EM fields

$$\vec{B}_5 = \vec{\nabla} \times \vec{A}_5 \quad \text{and} \quad \vec{E}_5 = -\vec{\nabla} A_0 - \partial_t \vec{A}_5$$



CONSISTENT CHIRAL KINETIC THEORY

[Gorbar, Miransky, Shovkovy, Sukhachov, Phys. Rev. Lett. **118**, 127601 (2017)]

[Gorbar, Miransky, Shovkovy, Sukhachov, Phys. Rev. B **95**, 115202 (2017)]

- Kinetic equation: [Son and Yamamoto, Phys. Rev. D **87**, 085016 (2013)]
[Stephanov and Yin, Phys. Rev. Lett. **109**, 162001 (2012)]

$$\frac{\partial f_\lambda}{\partial t} + \frac{\left[e\tilde{\mathbf{E}}_\lambda + \frac{e}{c}(\mathbf{v} \times \mathbf{B}_\lambda) + \frac{e^2}{c}(\tilde{\mathbf{E}}_\lambda \cdot \mathbf{B}_\lambda)\boldsymbol{\Omega}_\lambda \right] \cdot \nabla_{\mathbf{p}} f_\lambda}{1 + \frac{e}{c}(\mathbf{B}_\lambda \cdot \boldsymbol{\Omega}_\lambda)} + \frac{\left[\mathbf{v} + e(\tilde{\mathbf{E}}_\lambda \times \boldsymbol{\Omega}_\lambda) + \frac{e}{c}(\mathbf{v} \cdot \boldsymbol{\Omega}_\lambda)\mathbf{B}_\lambda \right] \cdot \nabla_{\mathbf{r}} f_\lambda}{1 + \frac{e}{c}(\mathbf{B}_\lambda \cdot \boldsymbol{\Omega}_\lambda)} = 0$$

where $\tilde{\mathbf{E}}_\lambda = \mathbf{E}_\lambda - (1/e)\nabla_{\mathbf{r}}\epsilon_{\mathbf{p}}$, $\mathbf{v} = \nabla_{\mathbf{p}}\epsilon_{\mathbf{p}}$, and

$$\epsilon_{\mathbf{p}} = v_{Fp} \left[1 - \frac{e}{c}(\mathbf{B}_\lambda \cdot \boldsymbol{\Omega}_\lambda) \right]$$

and $\boldsymbol{\Omega}_\lambda = \lambda\hbar \frac{\hat{\mathbf{p}}}{2p^2}$ is the Berry curvature

- The definitions of density and current are

$$\rho_\lambda = e \int \frac{d^3 p}{(2\pi\hbar)^3} \left[1 + \frac{e}{c} (\mathbf{B}_\lambda \cdot \boldsymbol{\Omega}_\lambda) \right] f_\lambda,$$

$$\mathbf{j}_\lambda = e \int \frac{d^3 p}{(2\pi\hbar)^3} \left[\mathbf{v} + \frac{e}{c} (\mathbf{v} \cdot \boldsymbol{\Omega}_\lambda) \mathbf{B}_\lambda + e (\tilde{\mathbf{E}}_\lambda \times \boldsymbol{\Omega}_\lambda) \right] f_\lambda$$

$$+ e \nabla \times \int \frac{d^3 p}{(2\pi\hbar)^3} f_\lambda \epsilon_{\mathbf{p}} \boldsymbol{\Omega}_\lambda,$$

They satisfy the following anomalous relations:

$$\frac{\partial \rho_5}{\partial t} + \nabla \cdot \mathbf{j}_5 = \frac{e^3}{2\pi^2 \hbar^2 c} \left[(\mathbf{E} \cdot \mathbf{B}) + (\mathbf{E}_5 \cdot \mathbf{B}_5) \right] \quad \checkmark$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{j} = \frac{e^3}{2\pi^2 \hbar^2 c} \left[(\mathbf{E} \cdot \mathbf{B}_5) + (\mathbf{E}_5 \cdot \mathbf{B}) \right] \quad \times$$

Consistent definition of current

- Bardeen-Zumino (Chern-Simons) term is needed,

$$\delta j^\mu = \frac{e^3}{4\pi^2 \hbar^2 c} \epsilon^{\mu\nu\rho\lambda} A_\nu^5 F_{\rho\lambda}$$

or

[Gorbar, Miransky, Shovkovy, Sukhachov, PRL **118**, 127601 (2017)]

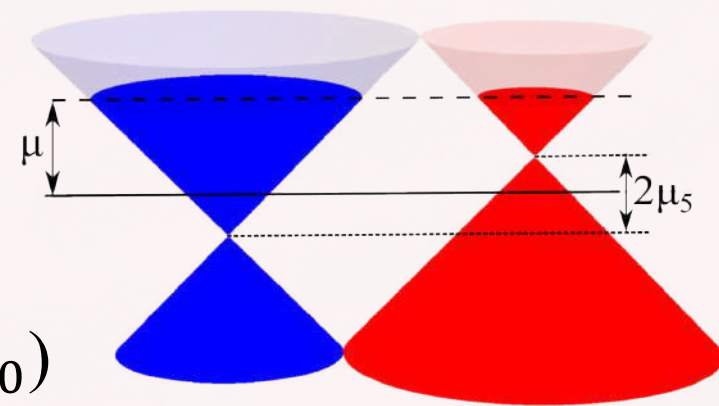
$$\delta \rho = -\frac{e^3}{2\pi^2 \hbar^2 c^2} (\mathbf{b} \cdot \mathbf{B})$$

$$\delta \mathbf{j} = -\frac{e^3}{2\pi^2 \hbar^2 c} b_0 \mathbf{B} + \frac{e^3}{2\pi^2 \hbar^2 c} [\mathbf{b} \times \mathbf{E}]$$

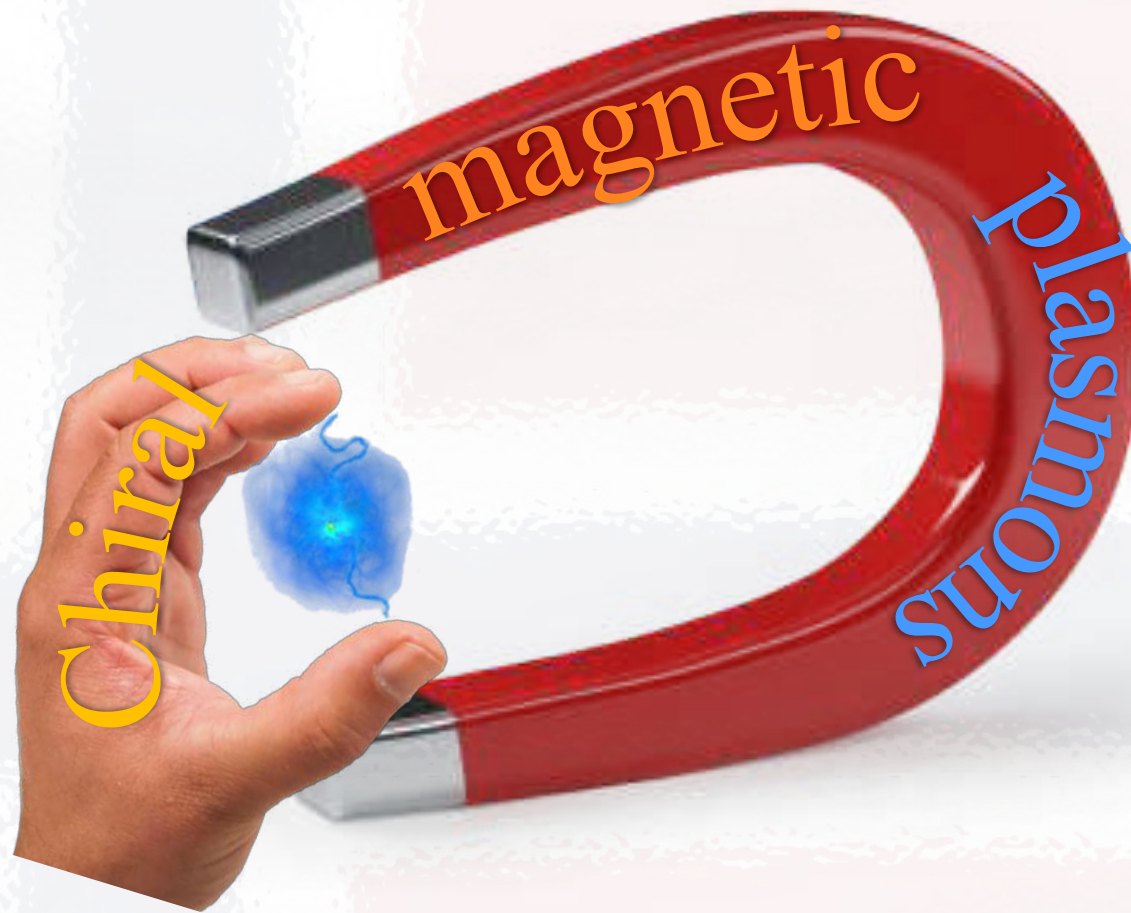
[Gorbar, Miransky, Shovkovy, Sukhachov, PRB **96**, 085130 (2017)]

- Its role and implications:

- Electric charge is conserved ($\partial_\mu J^\mu = 0$)
- Anomalous Hall effect is reproduced
- CME vanishes in equilibrium ($\mu_5 = -eb_0$)



- Any observable properties of Weyl materials directly sensitive to b_0 and \vec{b} ? Chiral anomaly? Topology?
- Potentially, there are many observable effects:
 - Anomalous Hall effect
 - Negative magnetoresistance
 - Strain/torsion induced CME
 - Quantum oscillations of the density of states
 - Strain/torsion dependent resistance
 - Unusual features of collective modes
 - Anomalous electric/chiral/thermal transport
 - Unusual features of nonlocal transport



CHIRAL MAGNETIC PLASMONS

[Gorbar, Miransky, Shovkovy, Sukhachov, Phys. Rev. Lett. **118**, 127601 (2017)]

[Gorbar, Miransky, Shovkovy, Sukhachov, Phys. Rev. B **95**, 115202 (2017)]

Chiral magnetic plasmons

Non-degenerate plasmon frequencies @ $k=0$:

$$\omega_l = \Omega_e, \quad \omega_{\text{tr}}^{\pm} = \Omega_e \sqrt{1 \pm \frac{\delta\Omega_e}{\Omega_e}}$$

where the Langmuir frequency is

$$\Omega_e \equiv \sqrt{\frac{4\alpha}{3\pi\hbar^2} \left(\mu^2 + \mu_5^2 + \frac{\pi^2 T^2}{3} \right)}$$

and

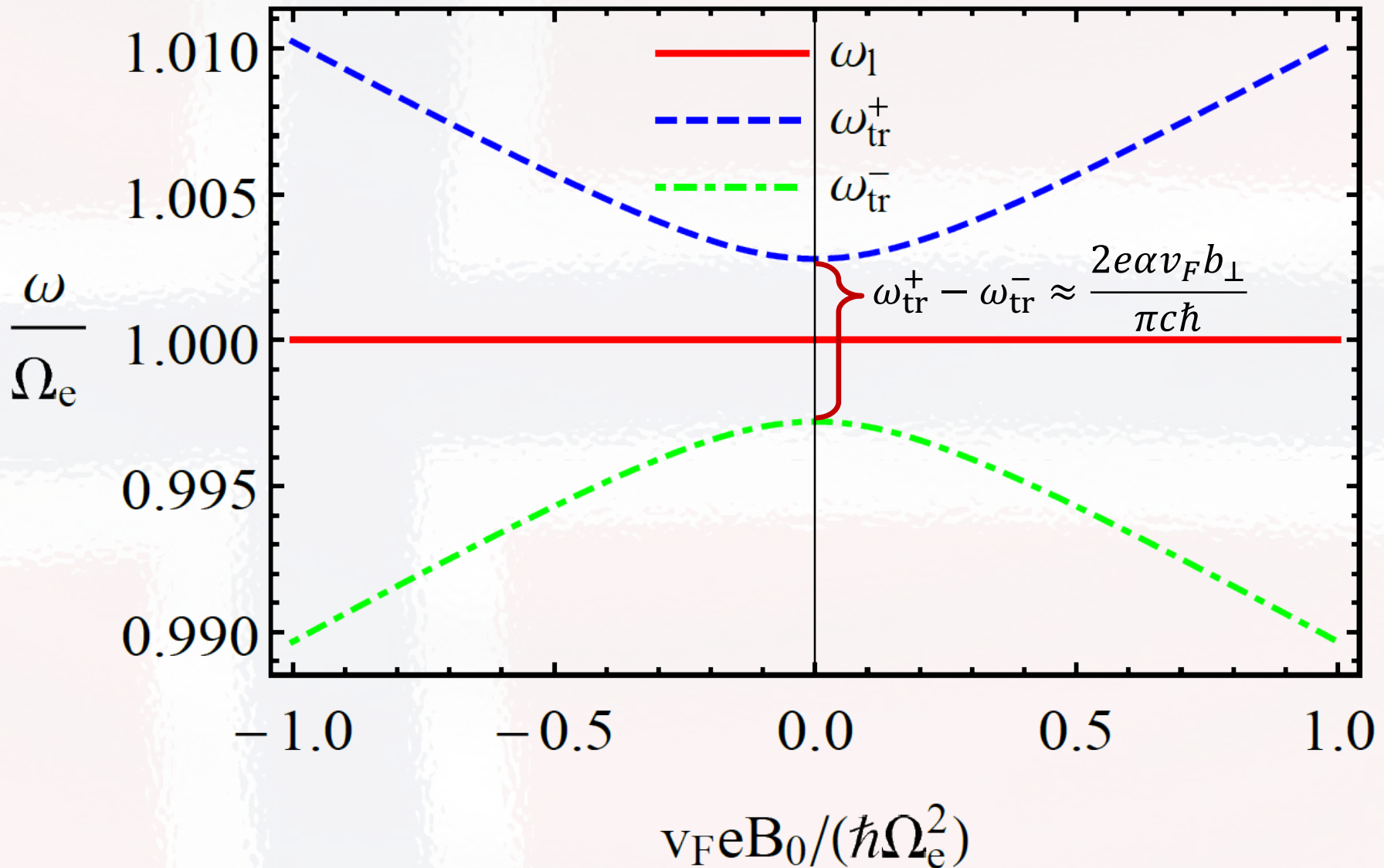
$$\delta\Omega_e = \frac{2e\alpha v_F}{3\pi c\hbar^2} \left\{ 9\hbar^2 b_{\perp}^2 + \left[\frac{2v_F}{\Omega_e^2} (B_0\mu + B_{0,5}\mu_5) - 3\hbar b_{\parallel} - \frac{v_F\hbar^2}{4T} \sum_{\lambda=\pm} B_{0,\lambda} F\left(\frac{\mu_{\lambda}}{T}\right) \right]^2 \right\}^{1/2}$$

[Gorbar, Miransky, Shovkovy, Sukhachov, Phys. Rev. Lett. **118**, 127601 (2017)]

[Gorbar, Miransky, Shovkovy, Sukhachov, Phys. Rev. B **95**, 115202 (2017)]

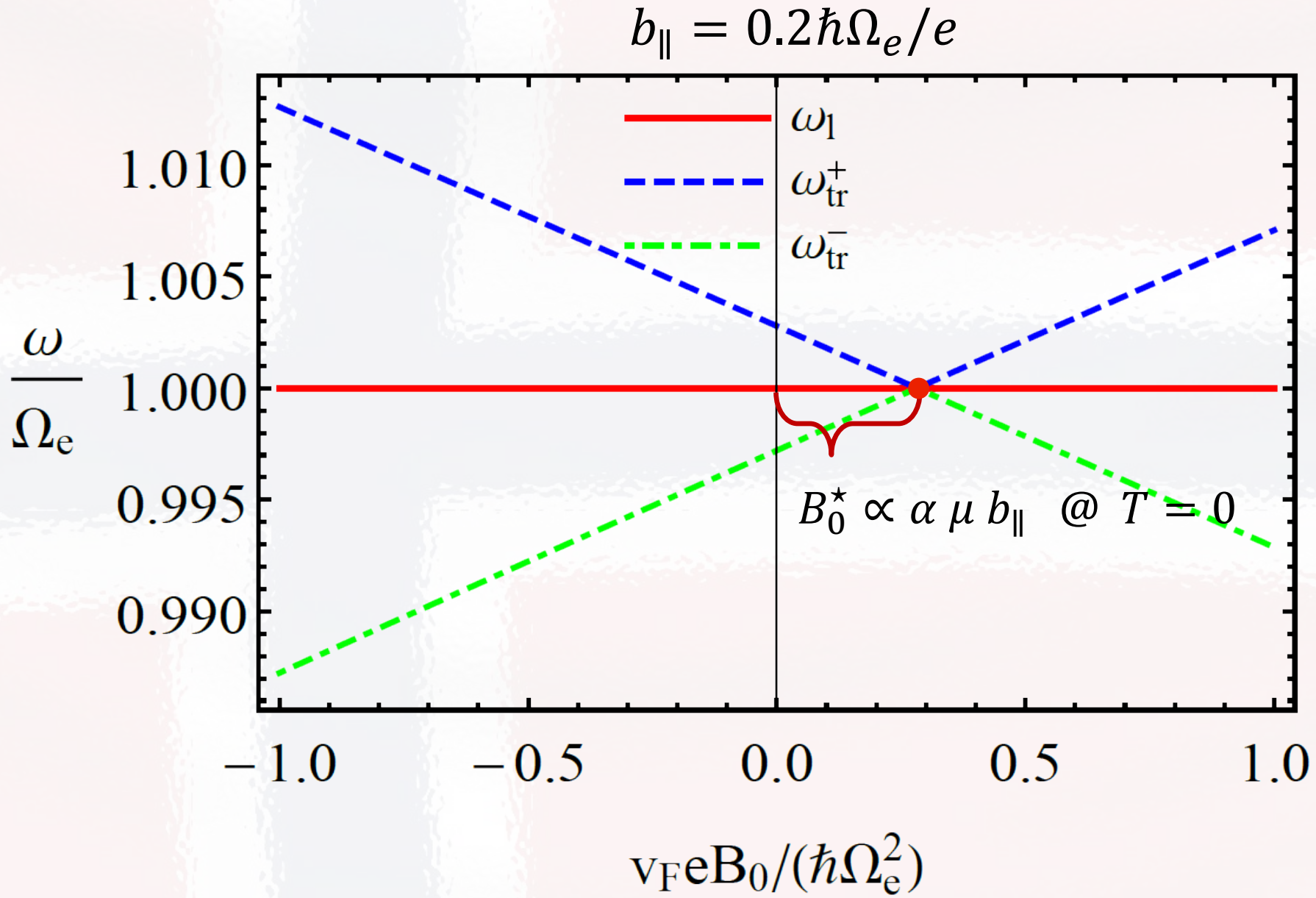
Plasmon frequencies, $\vec{B} \perp \vec{b}$

$$b_{\perp} = 0.2\hbar\Omega_e/e$$

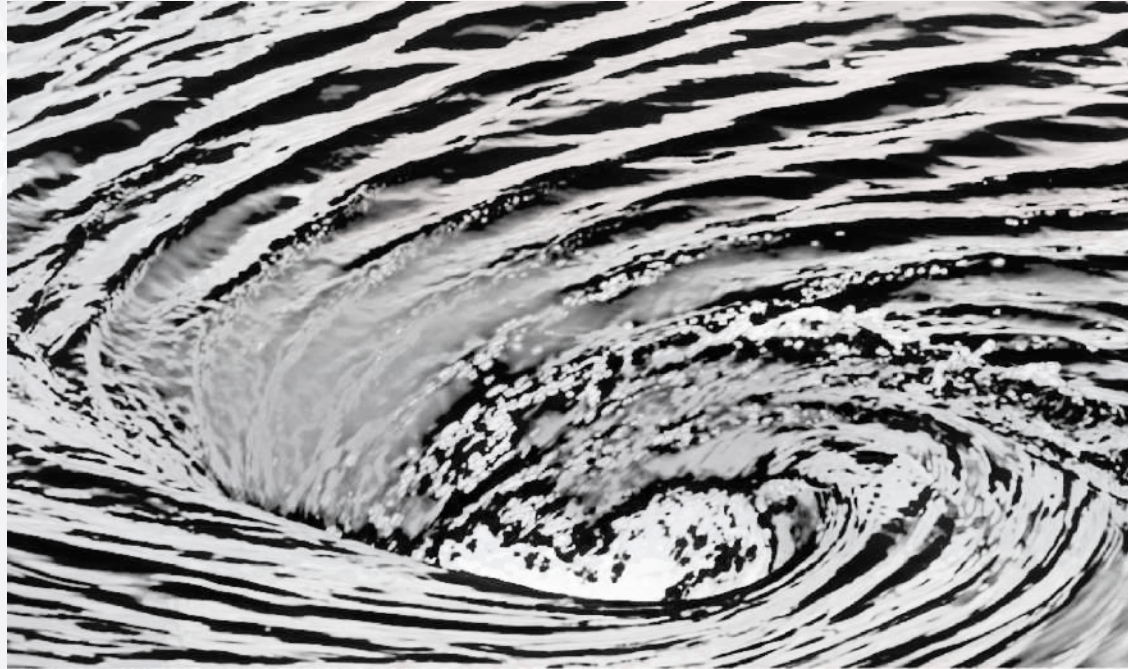


[Gorbar, Miransky, Shovkovy, Sukhachov, Phys. Rev. Lett. **118**, 127601 (2017)]

Plasmon frequencies, $\vec{B} \parallel \vec{b}$



[Gorbar, Miransky, Shovkovy, Sukhachov, Phys. Rev. Lett. **118**, 127601 (2017)]



ELECTRON TRANSPORT IN HYDRODYNAMIC REGIME

The Euler equation for electron fluid:

[Gurzhi, JETP 17, 521 (1963)]

$$\frac{1}{v_F} \partial_t \left(\frac{\epsilon + P}{v_F} \mathbf{u} + \sigma^{(\epsilon, B)} \mathbf{B} \right) = -en \left(\mathbf{E} + \frac{1}{c} [\mathbf{u} \times \mathbf{B}] \right) + \frac{\sigma^{(B)} (\mathbf{E} \cdot \mathbf{B})}{3v_F^2} \mathbf{u} - \frac{\epsilon + P}{\tau v_F^2} \mathbf{u} + O(\nabla_{\mathbf{r}})$$

[Gorbar, Miransky, Shovkovy, Sukhachov, PRB 97, 121105(R) (2018)]

The energy conservation

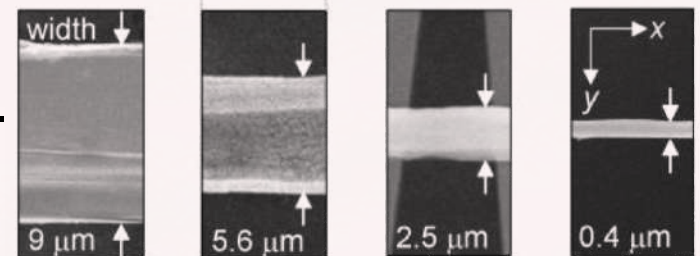
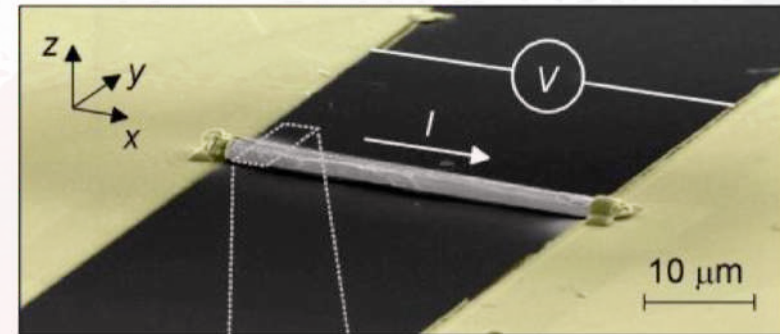
$$\partial_t \epsilon = -\mathbf{E} \cdot (en\mathbf{u} - \sigma^{(B)} \mathbf{B}) + O(\nabla_{\mathbf{r}})$$

+ Maxwell equations with the Chern-Simons currents

$$\rho_{CS} = -\frac{e^3 (\mathbf{b} \cdot \mathbf{B})}{2\pi^2 \hbar^2 c^2}$$

$$\mathbf{J}_{CS} = -\frac{e^3 b_0 \mathbf{B}}{2\pi^2 \hbar^2 c} + \frac{e^3 [\mathbf{b} \times \mathbf{E}]}{2\pi^2 \hbar^2 c}$$

Experimental evidence in tungsten diphosphide (WP_2) [Gooth et al., Nature Commun. 9, 4093 (2018)]



- Magneto-acoustic wave ($\rho = 0$):

$$\omega_{s,\pm} = -\frac{i}{2\tau} \pm \frac{i}{2\tau} \sqrt{1 - 4\tau^2 v_F^2 \frac{|\mathbf{k}|^2 w_0 - \sigma^{(\epsilon,u)} [2|\mathbf{k}|^2 B_0^2 - (\mathbf{k} \cdot \mathbf{B}_0)]}{3w_0}}$$

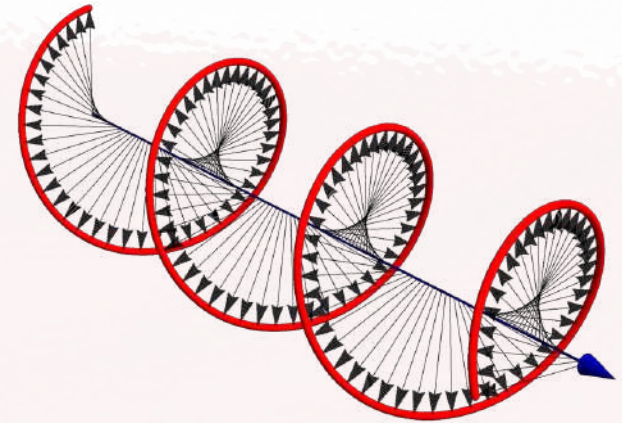
- *Gapped* chiral magnetic wave ($\rho = 0$):

$$\omega_{\text{gCMW},\pm} = \pm \frac{eB_0 \sqrt{3v_F^3 \left(4\pi e^2 T_0^2 + 3\varepsilon_e \hbar^3 v_F^3 k_{\parallel}^2\right)}}{2\pi^2 T_0^2 c \sqrt{\varepsilon_e \hbar}}$$

- Helicons ($\rho \neq 0$):

$$\omega_{h,\pm} \approx \mp \frac{ck_{\parallel}^2 B_0}{4\pi\mu_m \rho_0} - \frac{i}{\tau} \frac{c^2 k_{\parallel}^2 w_0}{4\pi\mu_m v_F^2 \rho_0^2} + O(k_{\parallel}^3)$$

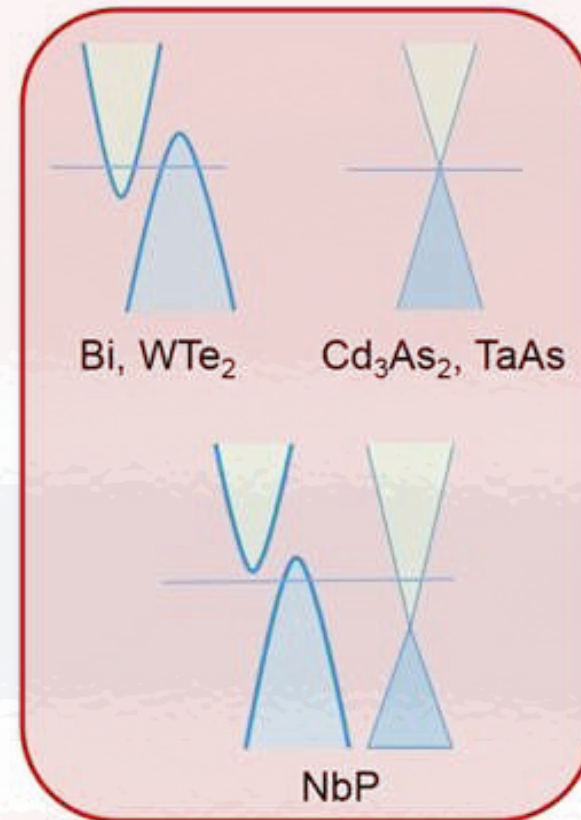
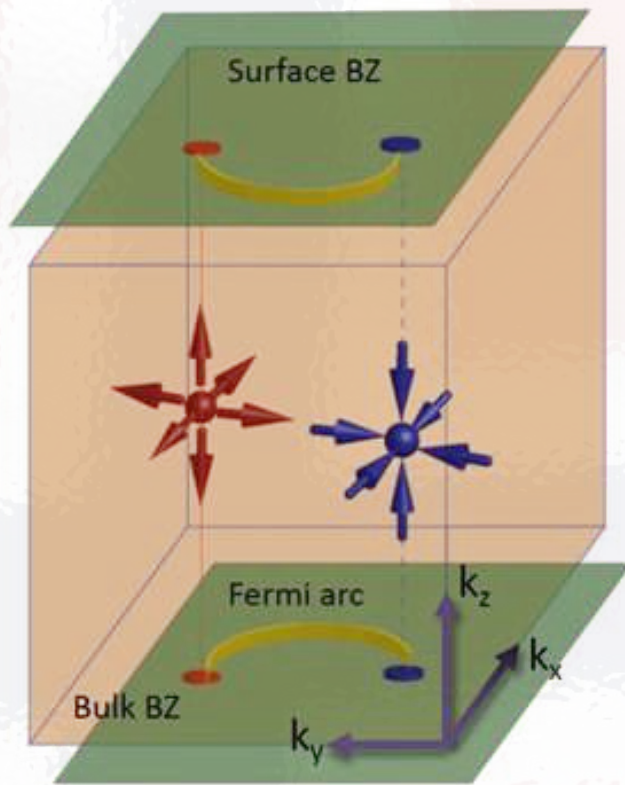
- New anomalous Hall waves at $\vec{b} \neq 0$, etc.



[Sukhachov, Gorbar, Shovkovy, Miransky, J. Phys. Cond. Matt. **30**, 275601 (2018)]

- **Neutral chiral plasma has chiral vortical waves**
 - Speeds depend on the direction @ $\mu_5 \neq 0$
[Gorbar, Rybalka, Shovkovy, Phys. Rev. D 95, 096010 (2017)]
(might be important for neutrino flow in supernovas/protoneutron stars)
- **Propagating (not overdamped) hydrodynamic modes in charged chiral plasma are** [Rybalka, Gorbar, Shovkovy, arXiv:1807.07608]
 - Sound and Alfven waves @ high temperature
 - Plasmons and helicons @ high density
- **Dynamical electromagnetism plays a crucial role**
 - no chiral magnetic waves, special chiral Alfven waves, or other exotic modes obtained in background-field approximation
 - chiral magnetic wave predictions for HIC should be revised

- Anomalous *macroscopic* effects are expected in many forms of chiral plasmas
- Experimental search for anomalous signatures in high-energy physics is extremely *difficult*
- Low-energy *chiral fermions* can be realized in Dirac/Weyl materials
- Fundamental anomalous physics could be tested in *table-top experiments*
- Chirality and anomaly could be valuable in *applied* research



EXTRAS

Dirac materials

- $\text{Bi}_{1-x}\text{Sb}_x$ alloy (at $x \approx 4\%$)

- Na_3Bi →

[Liu et al., Science 343, 864 (2014)]

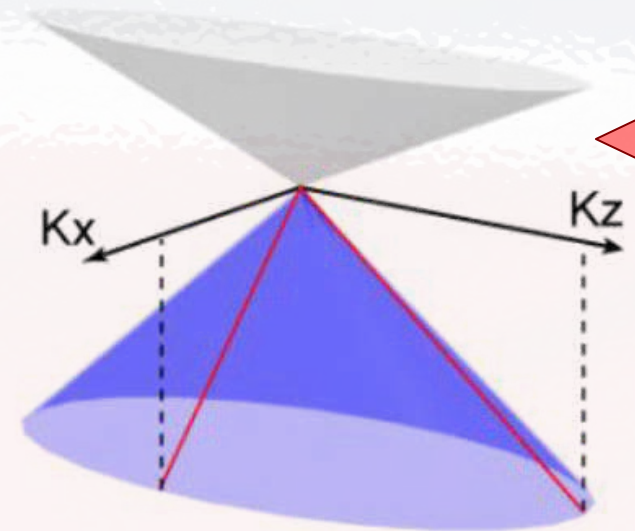
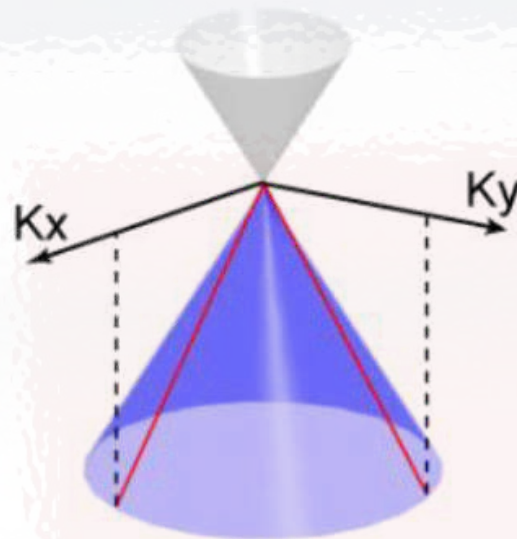
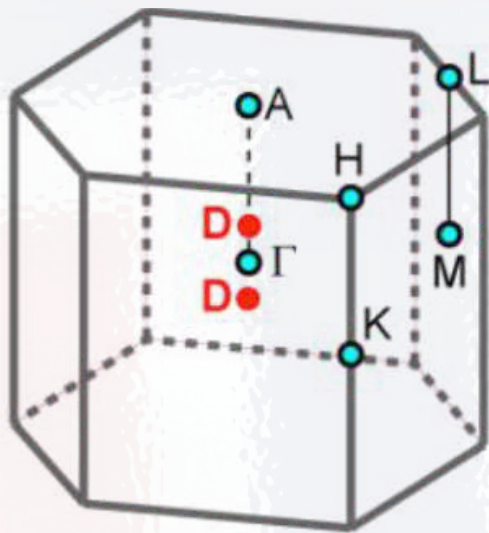
- Cd_3As_2 →

[Neupane et al., Nature Commun. 5, 3786 (2014)]

[Borisenko et al., Phys. Rev. Lett. 113, 027603 (2014)]

- ZrTe_5 →

[Li et al., Nature Physics 12, 550 (2016)]



$$E = \sqrt{v_x^2 k_x^2 + v_y^2 k_y^2 + v_z^2 k_z^2}, \quad v_x \approx v_y \approx 3.74 \times 10^5 \text{ m/s}, \quad v_z \approx 2.89 \times 10^4 \text{ m/s}$$

Weyl materials

- TaAs (tantalum arsenide) [S.-Y. Xu et al., Science 349, 613 (2015)]
[B. Q. Lv et al., Phys. Rev. X5, 031013 (2015)]
- NbAs (niobium arsenide) [S.-Y. Xu et al., Nature Physics 11, 748 (2015)]
- TaP (tantalum phosphide) [S.-Y. Xu et al., Science Adv. 1, 1501092 (2015)]
- NbP (niobium phosphide) [I. Belopolski et al. arXiv:1509.07465]
- WTe₂ (tungsten telluride) [F. Y. Bruno et al., Phys. Rev. B 94, 121112 (2016)]

