





# Anomalous physics of magnetized quark-gluon plasma Igor Shovkovy Arizona State University

[Miransky & Shovkovy, Phys. Rep. 576, 1 (2015)] [Kharzeev, Liao, Voloshin, Wang, Prog. Part. Nucl. Phys. 88, 1 (2016)]



#### Quark-gluon plasma

• Quark-gluon plasma (QGP) is a state of matter at high energy density (made of deconfined quarks and gluons)







#### Relativistic Matter

• Non-relativistic



- Particles move much slower than the speed of light
- Kinetic energies are much smaller than the rest energy

$$E_{\rm kin} \ll E_{\rm rest}$$
:  $E = c\sqrt{p^2 + m^2 c^2} \approx mc^2 + \frac{p^2}{2m}$ 

#### • Relativistic

- Particle velocities approach the speed of light
- Kinetic energies are comparable to, or larger than  $E_{\text{rest}}$

$$E_{\rm kin} \gtrsim E_{\rm rest}$$
:  $E = c\sqrt{p^2 + m^2 c^2} \approx cp$ 



• What happens when you squeeze matter to very high density?

**Pauli exclusion principle**: fermions cannot occupy same quantum states (they end up filling out all states from  $p_{\min} \approx 0$  to  $p_{\max} \propto \hbar n^{1/3}$ )





• What happens when you heat matter to very high temperature? (e.g., matter in heavy ion collisions)



Heat is equivalent to kinetic energy: average kinetic energy of particles is proportional to temperature:

 $p \approx k_B T/c \propto 200 \left(\frac{k_B T}{200 \text{ MeV}}\right) \text{MeV/c} (assuming <math>p \gg mc)$ 



#### Relativistic matter

- Early Universe (high temperature)
- Heavy-ion collisions (high temperature) [Kharzeev, Liao, Voloshin, Wang, Prog. Part. Nucl. Phys. 88, 1 (2016)]
- Super-dense matter in compact stars (high density)
- Ultra-relativistic jets from black holes (moderately high temperature and density)
- Dirac/Weyl (semi-)metals

[Gorbar, Miransky, Shovkovy, Sukhachov, *Electronic Properties of Dirac and Weyl Semimetals* (World Scientific, Singapore, 2021)]

• Other: cold atoms, superfluid <sup>3</sup>He-A, etc. [Volovik, JETP Lett. 105, 34 (2017)]





#### Chiral fermions

• Only *massless* Dirac fermions have a well-defined chirality  $(\gamma^5 \psi = \pm \psi)$ :



- *Massive* Dirac fermions have an *almost* well-defined chirality in the *ultrarelativistic* regime
  - High temperature:  $T \gg m$
  - High density:  $\mu \gg m$
- Chirality flip rate:  $\Gamma_{\rm flip} \propto \alpha^2 T (m/T)^2$



#### Anomalous chiral matter

- Relativistic matter made of chiral fermions may allow  $n_L \neq n_R$  on *macroscopic* scales
- The (collective) dynamics of  $n_R + n_L$  and  $n_R n_L$  is controlled by the continuity equations

$$\frac{\partial (n_R + n_L)}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0$$

$$\frac{\partial (n_R - n_L)}{\partial t} + \vec{\nabla} \cdot \vec{j}_5 = \frac{e^2 \vec{E} \cdot \vec{B}}{2\pi^2 c} - \Gamma_{\text{flip}}(n_R - n_L)$$

**Question**: Can chiral anomaly produce any *macroscopic* effects in ultra-relativistic matter?



# $\vec{B}$ and $\vec{\omega}$ in little Bangs

- Rotating & magnetized QGP created at RHIC/LHC
- Electromagnetic fields (Lienard-Wiechert potentials)



[Rafelski & Müller, PRL, 36, 517 (1976)], [Kharzeev et al., arXiv:0711.0950], [Skokov et al., arXiv:0907.1396], [Voronyuk et al., arXiv:1103.4239], [Bzdak &. Skokov, arXiv:1111.1949], [Deng & Huang, arXiv:1201.5108], ...

• Magnetic field estimate:  $B \sim 10^{18}$  to  $10^{19}$  G (~ 100 MeV)

• Vorticity estimate: [Adamczyk et al. (STAR), Nature 548, 62 (2017)]  $\omega \sim 9 \times 10^{21} s^{-1}$  (~ 10 MeV)





# CHIRAL SEPARATION EFFECT $\langle \vec{j}_5 \rangle = -\frac{e\vec{B}}{2\pi^2}\mu$

[Vilenkin, Phys. Rev. D 22 (1980) 3067] [Metlitski & Zhitnitsky, Phys. Rev. D 72, 045011 (2005)] [Newman & Son, Phys. Rev. D 73 (2006) 045006]

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#### Landau levels

Landau energy levels at m = 0 $E_n^{\pm} = \pm \sqrt{2n|eB|} + p_z^2$ where  $n = k + \frac{1}{2} + \operatorname{sgn}(eB)s_z$ orbital spin Lowest Landau level is spin polarized  $E_0^{\pm} = \pm p_z$   $(k = 0, s_z = -\frac{1}{2})$ Density of states at *E*=0:  $\frac{dn}{dE}\Big|_{E=0} = \frac{|eB|}{2\pi} \frac{1}{2\pi} = \frac{|eB|}{4\pi^2}$ Higher Landau levels  $(n \ge 1)$  are twice as degenerate: (i) k = n &  $s = -\frac{1}{2}$ (ii) k = n - 1 &  $s = +\frac{1}{2}$ 





#### Landau spectrum & $\mu \neq 0$





- Spin polarized LLL is chirally asymmetric
  - states with  $p_3 < 0$  (and  $s = \downarrow$ ) are R-handed
  - states with  $p_3 > 0$  (and  $s = \downarrow$ ) are L-handed
  - i.e., a nonzero **axial** current is induced  $\langle \vec{j}_5 \rangle = -tr[\vec{\gamma}\gamma^5 S(x,x)] = -\frac{e\vec{B}}{2\pi^2}\mu$







### CHIRAL MAGNETIC EFFECT

$$\langle \vec{j} \rangle = \frac{e^2 \vec{B}}{2\pi^2} \mu_5$$

[Fukushima, Kharzeev, Warringa, Phys. Rev. D 78, 074033 (2008)]

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#### Chiral Magnetic Effect ( $\mu_5 \neq 0$ )

Topological fluctuations could induce *transient* state with a nonzero chiral charge ( $\mu_5 \neq 0$ )

Spin polarized LLL ( $s=\downarrow$  for particles of a *negative* charge):

- Some R-handed states  $(p_3 < 0$ and  $E < \mu_5)$  are occupied
- Some L-handed holes (p<sub>3</sub> < 0 and |E| < μ<sub>5</sub>) are empty



CME current:

$$\langle \vec{j} \rangle = -tr[\vec{\gamma}S(x,x)] = \frac{e^2\vec{B}}{2\pi^2}\mu_5$$

[Fukushima, Kharzeev, Warringa, Phys. Rev. D 78, 074033 (2008)]



### Partially filled LLL ( $\mu_5 \neq 0$ )

• Spin polarized LLL is chirally asymmetric

- states with  $p_3 < 0$  (and s= $\downarrow$ ) are R-handed quarks

- states with  $p_3 > 0$  (and s= $\downarrow$ ) are L-handed **antiquarks**
- i.e., a nonzero electric current is induced





#### **HEAVY-ION COLLISIONS**

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#### ASJ CME in heavy ion collisions?

• Chiral charge can be produced by topological configurations in QCD

$$\frac{d(N_R - N_L)}{dt} = -\frac{g^2 N_f}{16\pi^2} \int d^3x \ F_a^{\mu\nu} \tilde{F}_{\mu\nu}^a$$

• A random fluctuation with nonzero chirality could result in

$$N_R - N_L \neq 0 \implies \mu_5 \neq 0$$

• This should lead to an electric current  $\langle \vec{j} \rangle = \frac{e^2 \vec{B}}{2 \pi^2} \mu_5$ 



#### Dipole CME

• Dipole pattern of electric currents (or charge correlations) in heavy ion collisions



[Kharzeev, McLerran, Warringa, Nucl. Phys. A **803**, 227 (2008)] [Fukushima, Kharzeev, Warringa, Phys. Rev. D **78**, 074033 (2008)]





#### Correlations of same & opposite charge particles:

[Abelev et al. (STAR), PRL **103**, 251601 (2009)] [Abelev et al. (STAR), PRC **81**, 054908 (2010)] [Abelev et al. (ALICE), PRL **110**, 012301 (2013)] [Adamczyk et al. (STAR), PRC **88**, 064911 (2013)]

$$\begin{cases} \langle \cos(\varphi_{\alpha}^{+} + \varphi_{\beta}^{+} - 2\Psi_{\rm RP}) \rangle \\ \langle \cos(\varphi_{\alpha}^{+} + \varphi_{\beta}^{+} - 2\Psi_{\rm RP}) \rangle \end{cases}$$

#### Large background effects!

[Belmont & Nagle, PRC 96, 024901 (2017)], [ALICE Collaboration, Phys. Lett. B777, 151 (2018)]

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#### CHIRAL MAGNETIC WAVE: NAÏVE APPROACH

[Gorbar, Miransky, Shovkovy, Phys. Rev. D 83, 085003 (2011)] [Burnier, Kharzeev, Liao, Yee, Phys. Rev. Lett. 107 (2011) 052303]

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#### Chiral Magnetic Wave

• Nonzero charge density (a)  $B \neq 0 \rightarrow CMW$ 



[Gorbar, Miransky, Shovkovy, Phys. Rev. D 83, 085003 (2011)]

• Back-to-back electric currents, or quadrupole charge correlations (i.e., difference in elliptic flows of in  $\pi^+$  and  $\pi^-$ )

$$\frac{dN_{\pm}}{d\phi} \approx \bar{N}_{\pm} [1 + 2v_2 \cos(2\phi) \mp A_{\pm} r \cos(2\phi)]$$



where  $A_{\pm}$  is the charge asymmetry

[Burnier, Kharzeev, Liao, Yee, Phys. Rev. Lett. 107 (2011) 052303]

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#### Dispersion of CMW $(\mathbf{k} \parallel \mathbf{B})$

• Simple model  $(\delta n, \delta n_5 \sim e^{-i\omega t + ikz})$ :

$$k_0 \delta n - \frac{eB}{2\pi^2 \chi_5} k \delta n_5 = 0$$
$$k_0 \delta n_5 - \frac{eB}{2\pi^2 \chi} k \delta n = 0$$

where  $\chi_5 \simeq \chi = \partial n / \partial \mu \simeq T^2 / 3$ 

• The linear dispersion of the CMW mode:

$$k_0 \simeq \pm \frac{eB}{2\pi^2 \chi} k$$

• This is a gapless mode with speed  $v \propto eB/T^2$ 

#### **ASJ** CMW: Experimental evidence



[Ke (for STAR) J. Phys. Conf. Series **389**, 012035 (2012)] [Adamczyk et al. (STAR), Phys. Rev. Lett. **114**, 252302 (2015)] [Adam et al. (ALICE), Phys. Rev. C **93**, 044903 (2016)]

Higher harmonics of particle correlations are problematic...

#### Background effects may dominate over the signal!

[CMS Collaboration, arXiv:1708.08901]

In fact, the chiral magnetic wave might be overdapmed...

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## CHIRAL MAGNETIC WAVE: MORE RIGOROUS

[Rybalka, Gorbar, Shovkovy, arXiv:1807.07608, Phys. Rev. D 99, 016017 (2019)] [Shovkovy, Rybalka, Gorbar, arXiv:1811.10635, PoS (Confinement2018) 029]

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• The dispersion of the CMW mode:

$$k_0^{(\pm)} = -i\frac{\sigma_E}{2} \pm i\frac{\sigma_E}{2}\sqrt{1 - \left(\frac{3eB}{\pi^2 T^2 \sigma_E}\right)^2 \left(k^2 + \frac{e^2 T^2}{3}\right) - i\frac{\tau}{3}k^2}$$

• This is a completely diffusive mode when

$$\frac{3eB}{\pi^2 T^2 \sigma_E} \sqrt{k^2 + \frac{e^2 T^2}{3}} < 1$$



### CMW in QGP

• Model with two light u- and d-quarks:

$$k_0 \delta n_f - \frac{eq_f Bk}{2\pi^2 \chi_{f,5}} \delta n_{f,5} + iD_f k^2 \delta n_f - \frac{1}{eq_f} \sigma_{E,f} k \delta E_z = 0$$
  

$$k_0 \delta n_{f,5} - \frac{eq_f Bk}{2\pi^2 \chi_f} \delta n_f + iD_f k^2 \delta n_{f,5} - i\frac{e^2 q_f^2}{2\pi^2} B \delta E_z = 0$$
  

$$k \delta E_z + ie \sum_f q_f \delta n_f = 0$$
  
where  $f = u, d$ , and  $q_u = \frac{2}{3}, q_d = -\frac{1}{3}$ 

 $\chi_f$ ,  $D_f$ , and  $\sigma_{E,f}$  are susceptibilities, diffusion coefficients and electrical conductivities for each quark flavor, respectively



#### Non-perturbative regime

Near-critical strongly coupled quark-gluon plasma

$$\sigma_{E} = \sum_{f} \sigma_{E,f} = c_{\sigma} C_{em}^{\ell} T$$

$$\chi_{f} = c_{\chi} \chi_{f}^{(SB)}$$

$$C_{em}^{\ell} = (5/9) 4\pi \alpha_{em} \approx 0.051$$

$$D_{f} = \frac{c_{D}}{2\pi T}$$

Lattice data [Aarts, et. al. JHEP 1502, 186 (2015)]

	$c_{\sigma}$	c <sub>x</sub>	c <sub>D</sub>
T=200 MeV	0.111	0.804	0.758
T=235 MeV	0.214	0.885	1.394
T=350 MeV	0.316	0.871	1.826



#### Results

Two sets of modes  $k_{0,n}^{(\pm)} = \pm E_n(k) - i\Gamma_n(k)$ CMW is completely diffusive at small *eB* & *k*:





#### Summary

- Chiral anomalous effects can have observable signatures in (quasi-)relativistic matter
- Dynamical electromagnetism is important for CMW
  - Electrical conductivity screens charge fluctuations
  - Anomaly makes wave gapped
  - Charge diffusion is not negligible in finite-size systems
- Chiral magnetic wave is likely to be badly overdamped
- Empirical data will have to resolve the controversy
- Similar effects can be tested/observed in Dirac/Weyl semimetals, trapped cold atoms, superfluid helium, etc.